



Radiation-Driven Dynamic Target Response for Dissimilar Material Jetting and for Debris Effects in Partitioned Pipes

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Dissimilar Material Jetting on NOVA and Z

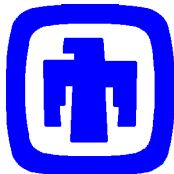
**T. A. Mehlhorn, T. A. Haill,
K. G. Budge, T. G. Trucano,
K. R. Cochran, J. J. MacFarlane**

Debris Effects in Partitioned Pipes

**M. D. Furnish, C. A. Hall,
J. R. Asay, T. G. Trucano,
K. G. Budge**

SAND2001-0183P

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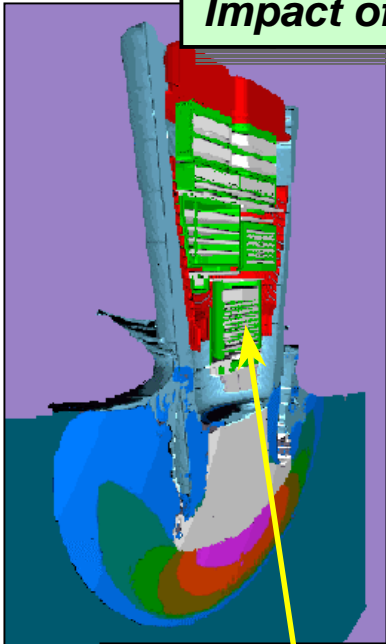


ALEGRA has unique capabilities for addressing SBSS program issues as well as HEDP problems.

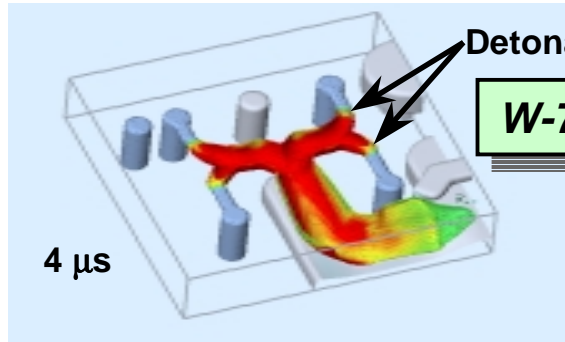
Z-pinch and Target Theory Department



Impact of Tactical Warhead



Critical components must survive the impact-induced shock loading

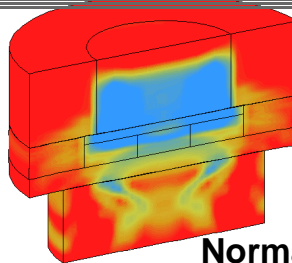


Detonation fronts

W-76 Fireset

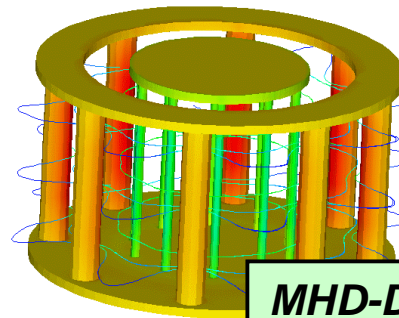
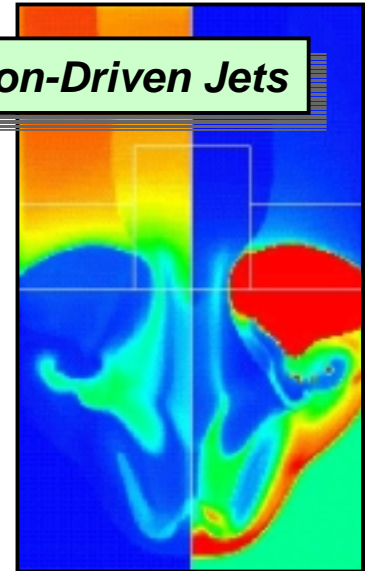
4 μ s

Dynamic Response of NG Power Supply



Normal stress (axial)

Radiation-Driven Jets



MHD-Driven Z Pinch

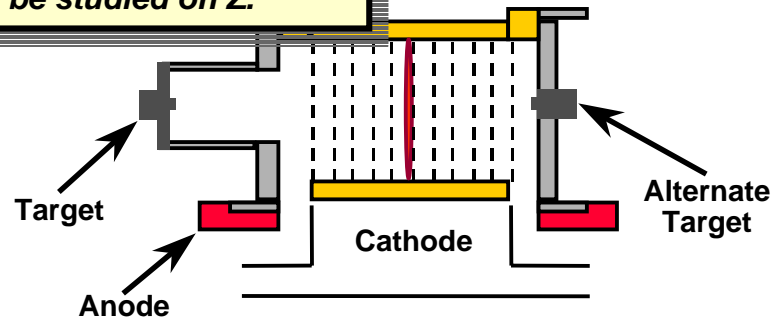
... the code combines solid dynamics, fracture, HE, etc., with high-energy features such as MHD and radiation transport.



Dissimilar Material Jetting on NOVA and Z

**R. J. Lawrence,
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*Similar targets scaled up
by an order of magnitude
can be studied on Z.*



*Models have been validated by
comparison with experimental
results from NOVA.*

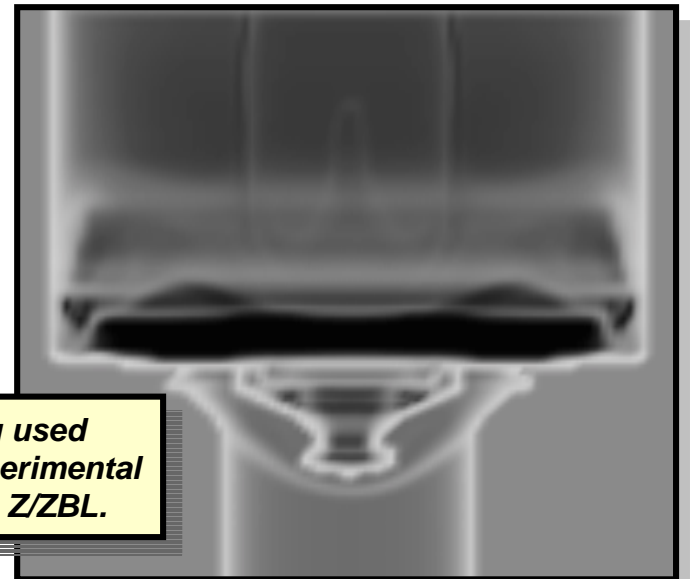
8 Nova laser beams
(16 TW total power)

3-mm Scale-1 hohlraum

X-ray camera

Ti or Fe backlighter

*ALEGRA and SPECT3D are being used
to calculate and visualize the experimental
results of RadJet experiments on Z/ZBL.*



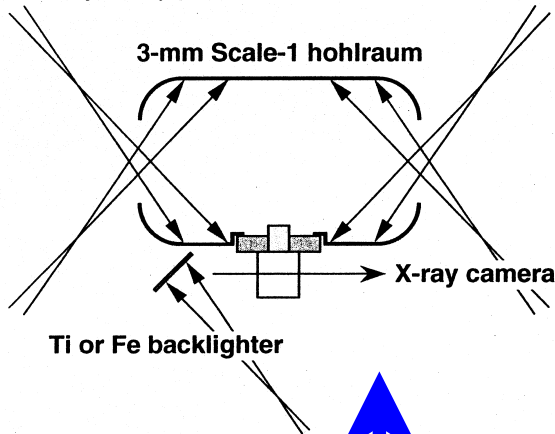


The NOVA experiments used radiation from a short-pulse, high-power, laser-driven hohlraum.

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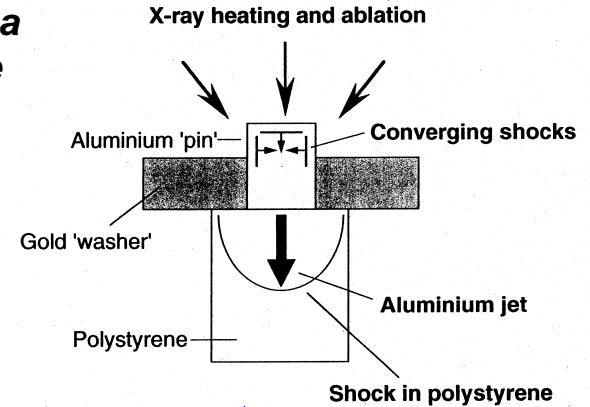
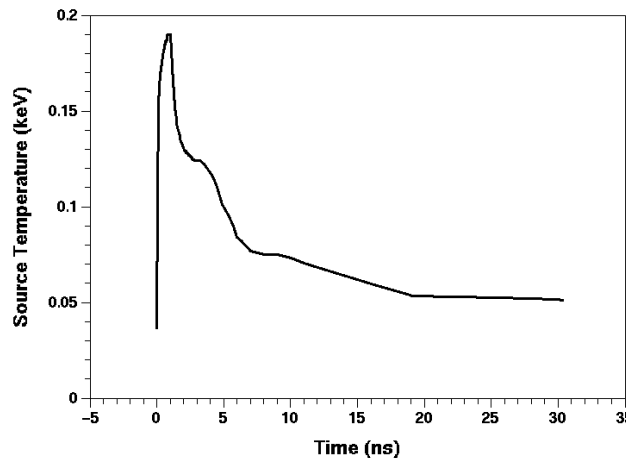


8 Nova laser beams
(16 TW total power)



The experiment used a laser-driven NOVA hohlraum to expose the sample to a short high-intensity radiation load. The sample response was observed with an x-ray backlighter.

The radiation drive was a blackbody temperature history peaking at ~190 eV with a FWHM pulse width of ~5 ns.



The target configuration consisted of a 150- μ m-long aluminum "pin" in a 50- μ m-thick gold "washer," which was backed with a 380- μ m-diameter polystyrene block.



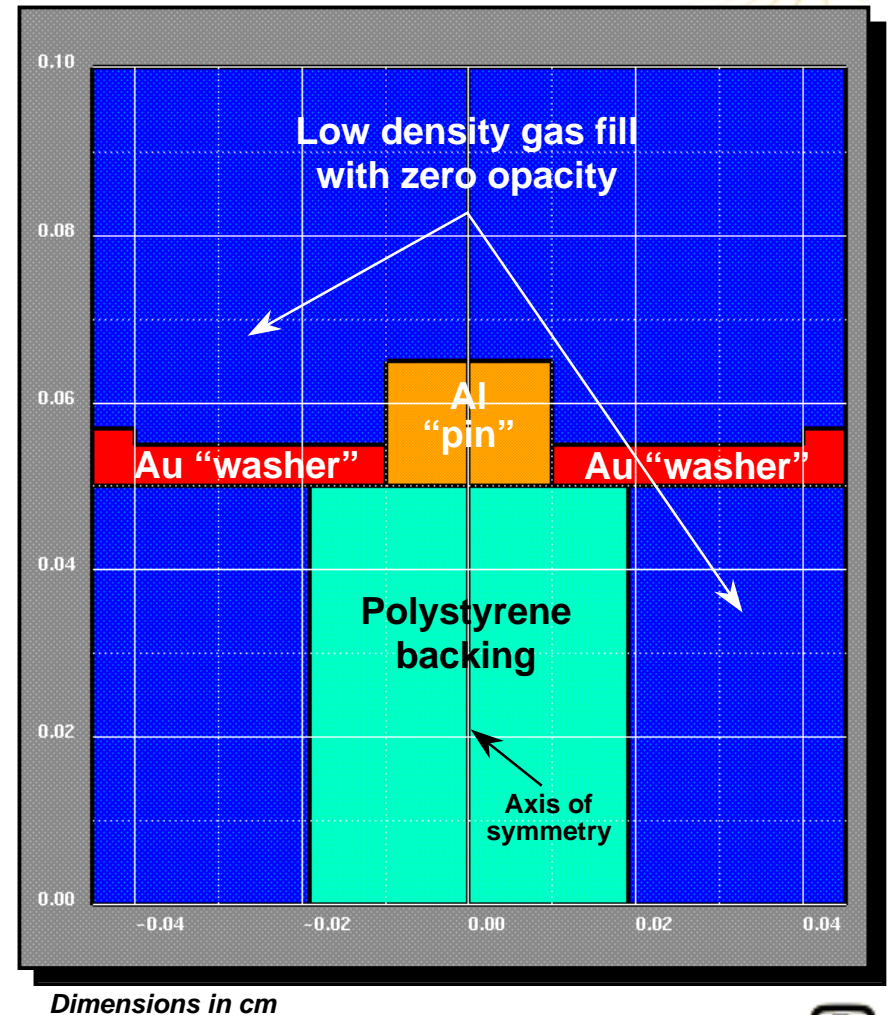


The configuration used for the ALEGRA calculations employed an Al “pin” mounted in an Au “washer.”

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- *For the NOVA problem, the thickness of the aluminum “pin” was $150\text{ }\mu\text{m}$, the gold washer thickness was $50\text{ }\mu\text{m}$, and the polystyrene backing had a diameter of $380\text{ }\mu\text{m}$.*
- *We used a 2-D cylindrical Eulerian mesh with: 1) 4,500 elements ($10\text{-}\mu\text{m}$ resolution); and 2) 18,000 elements ($5\text{-}\mu\text{m}$ resolution).*
- *The radiation, incident from the top, was treated with single-group, SN_1 radiation transport, with radiation pressure disabled.*



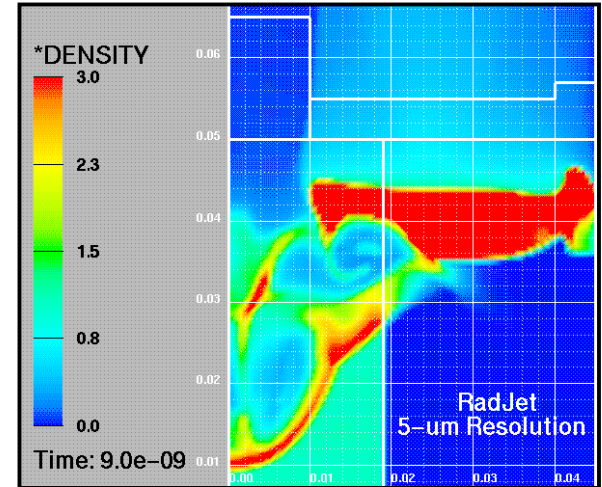
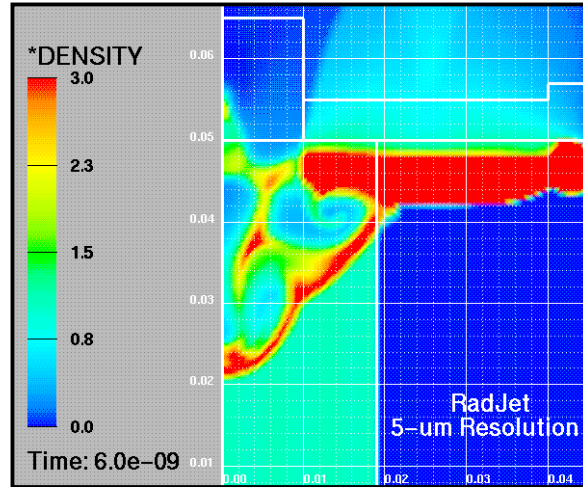
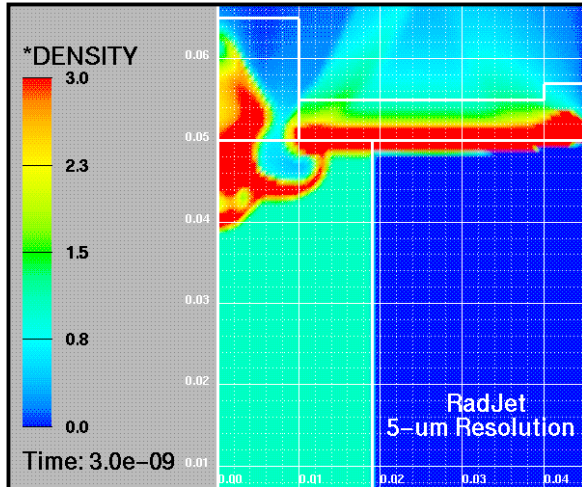
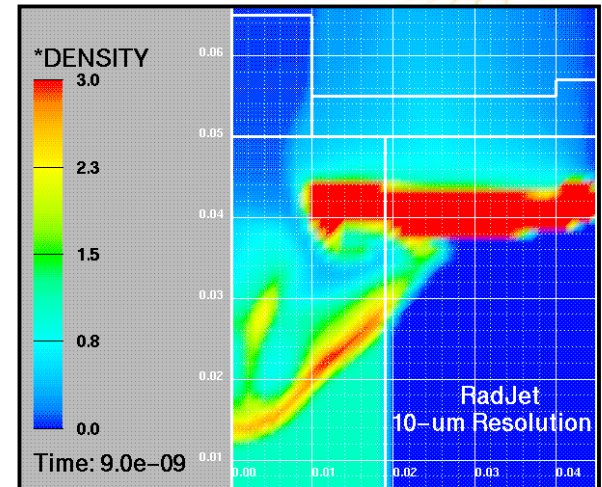
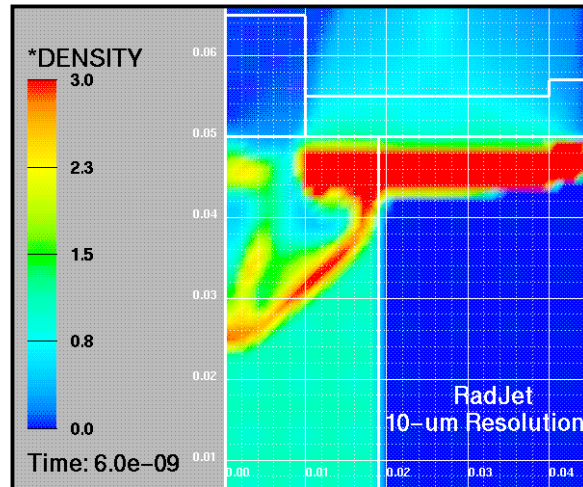
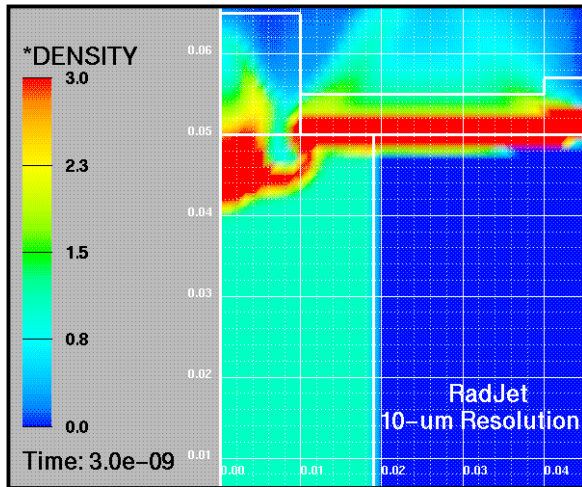


For NOVA, fine resolution calculations show more detail and slightly faster on-axis jet motion.

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Coarse resolution calculation →



Fine resolution calculation →



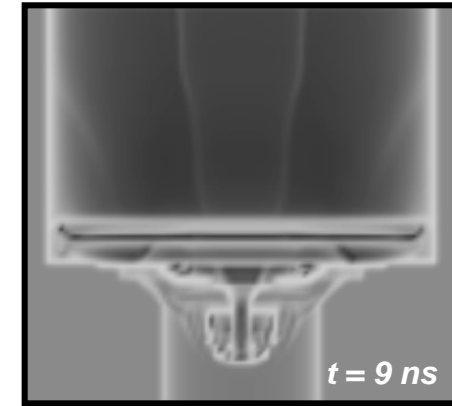
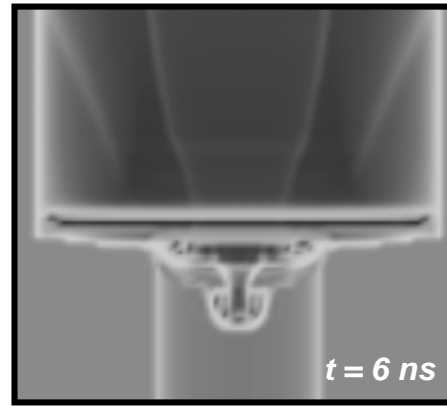
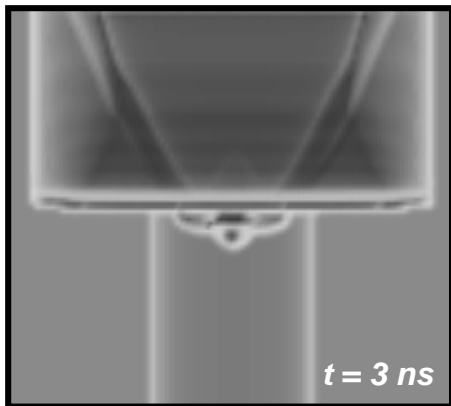
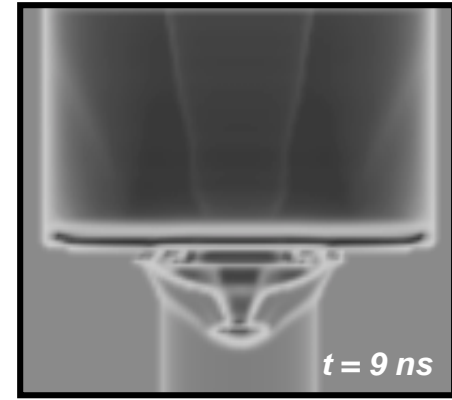
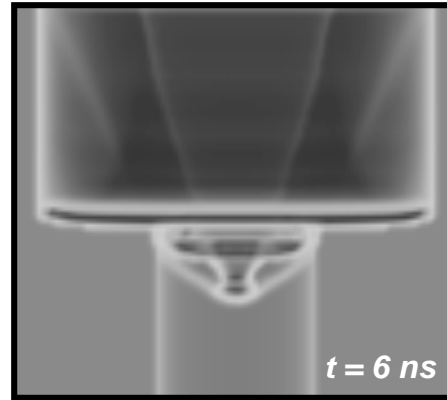
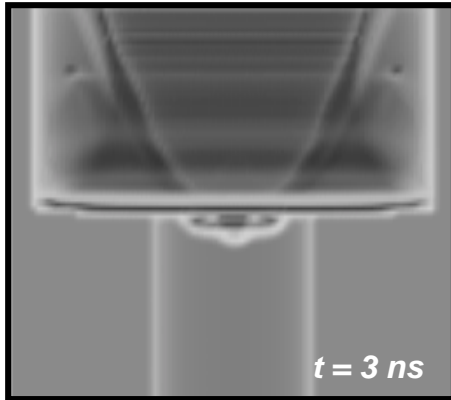


SPECT3D produces simulations of detector images from ALEGRA rad/hydro output.

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Coarse resolution calculation →



Fine resolution calculation →





For this configuration the ALEGRA results are consistent with other codes and the experiment.

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Spatial characterization of aluminum jet (Revised configuration – coarse mesh):

[Axial position of leading edge (μm)]

Code	ALEGRA (Eulerian)	PETRA (Eulerian)	CALE (ALE)	RAGE (AMR)	Experiment (Estimated)
Time = 6 ns	265	245	300	280	~260
Time = 9 ns	380	345	405	380	300+
Time = 12 ns	460	–	–	–	–

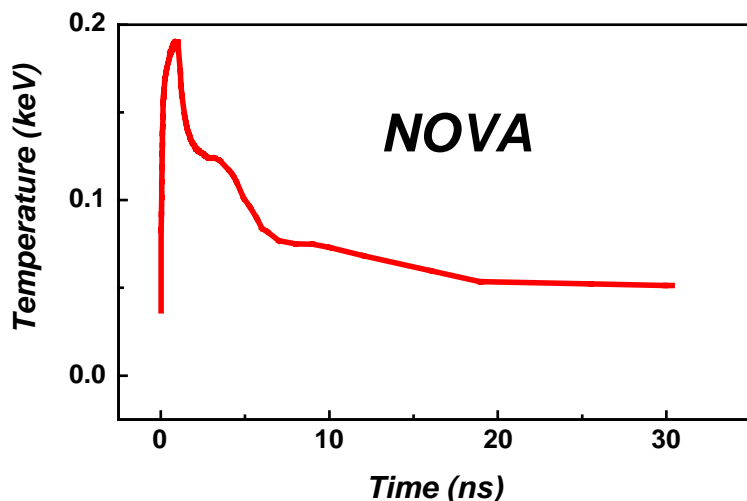
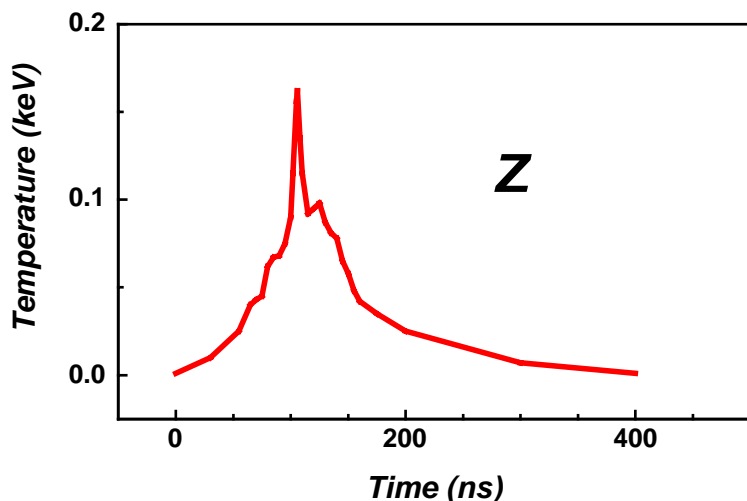
- *At a computational time of 6 ns, ALEGRA predicts the on-axis jet location within about 2% of the estimated experimental result; this result is also consistent with the other computational efforts.*
- *At a time of 9 ns the predicted axial location of the jet is somewhat over 20% greater than the estimated experimental measurement; but as with the earlier time, it agrees very closely with the average of the other code results.*





We are now studying the scaling of these RadJet experiments from NOVA to Z.

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Typical Source Characteristics

Peak temperature:

NOVA ~190 eV

Z ~162 eV

FWHM pulse width:

NOVA ~5 ns

Z ~50 ns

Peak power:

$$P_{\text{MAX}}(\text{Z})/P_{\text{MAX}}(\text{NOVA}) \approx 1/2$$

Total energy fluence:

$$\Phi_{\text{TOT}}(\text{Z})/\Phi_{\text{TOT}}(\text{NOVA}) \approx 3$$

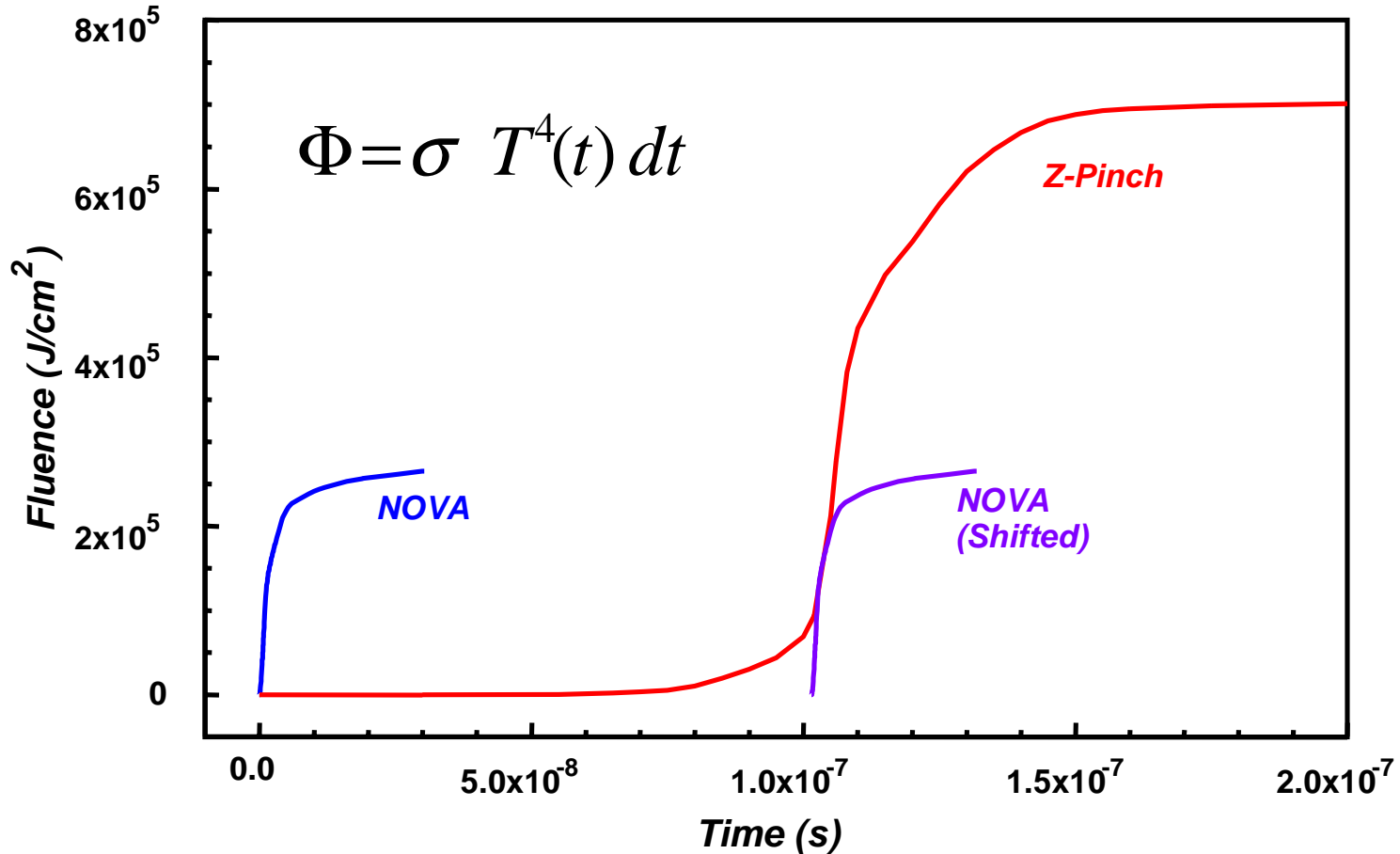
- *Similar mechanical behavior should be obtained by scaling the physical dimensions by about a factor of ten.*
- *However, the radiation transport will not scale in a similar fashion.*
- *Source for Z can be modified.*





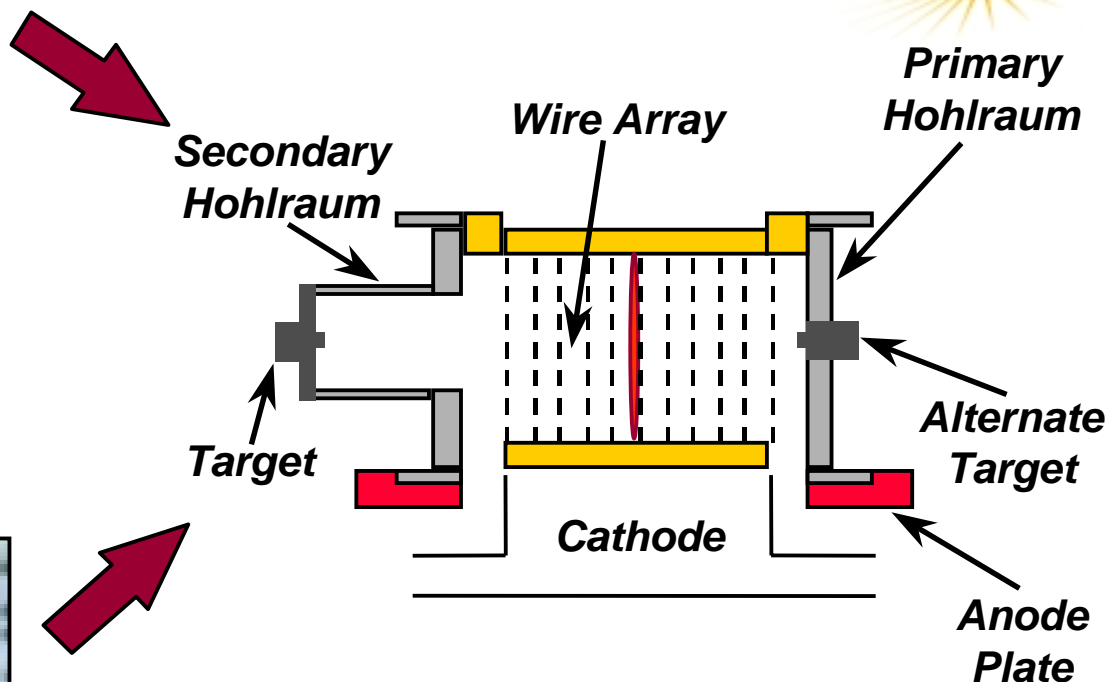
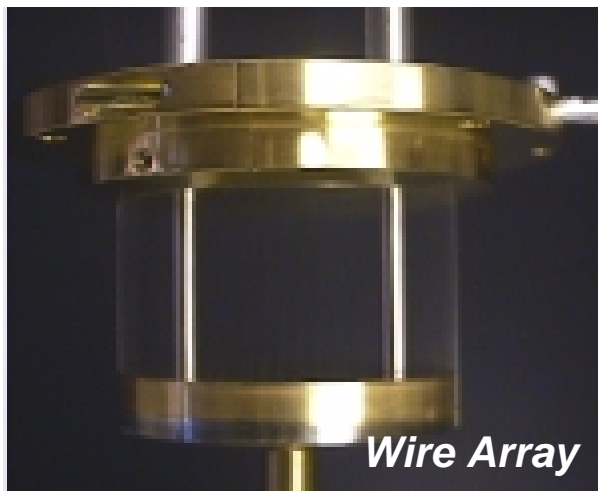
Integrating the temperature curves allows the fluences to be compared.

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The Z experiments would use a longer pulse, but a higher energy Z-pinch-driven hohlraum.

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- The target could be placed in either the primary or a secondary hohlraum.
- Principal diagnostics would involve the Z-Beamlet Backlighter that is presently under construction.



There are several points that should be noted with regard to the calculations.

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- The quoted half-max pulse widths are only approximate, but lead to about a factor of ten difference in characteristic response times.
- In these calculations the physical dimensions are scaled by exactly a factor of ten for the two cases.
- Because the radiation transport phenomena (e.g., opacities) do not scale in the same manner as the hydrodynamic behavior, the total response will not be directly homologous.
- The calculations were run with ALEGRA, using 10- μm resolution for the NOVA case and 100- μm resolution for the Z configuration.
- Because of the initial slow rise for the radiation drive from Z, the times cannot be shifted in a directly proportional fashion; the comparison plots were chosen for similar stages in the evolution of the response.





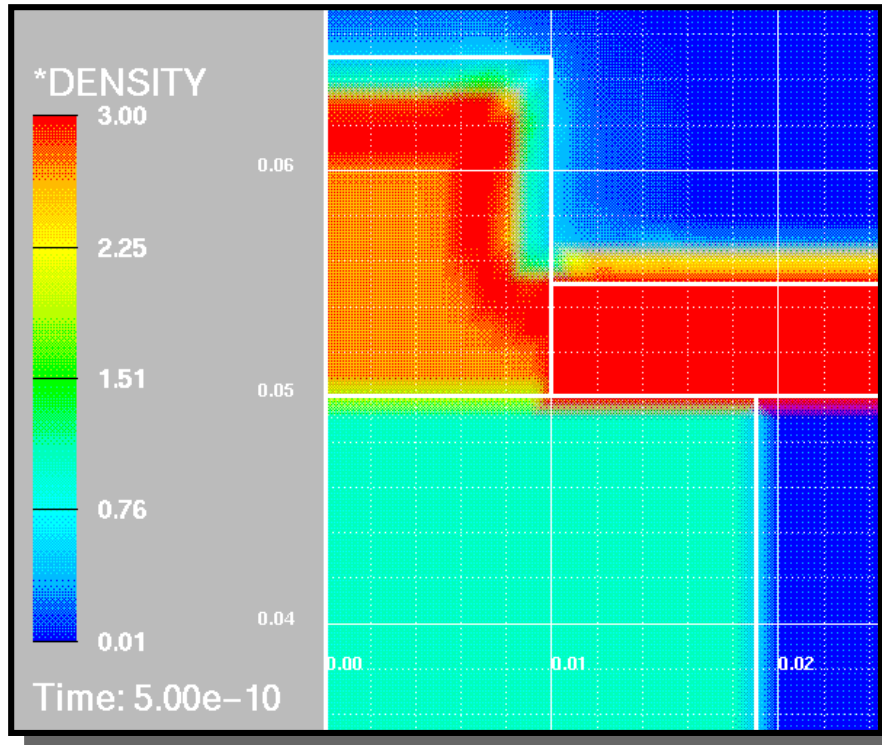
The RadJet problem at very early times, with the shock part way through the “pin” . . .

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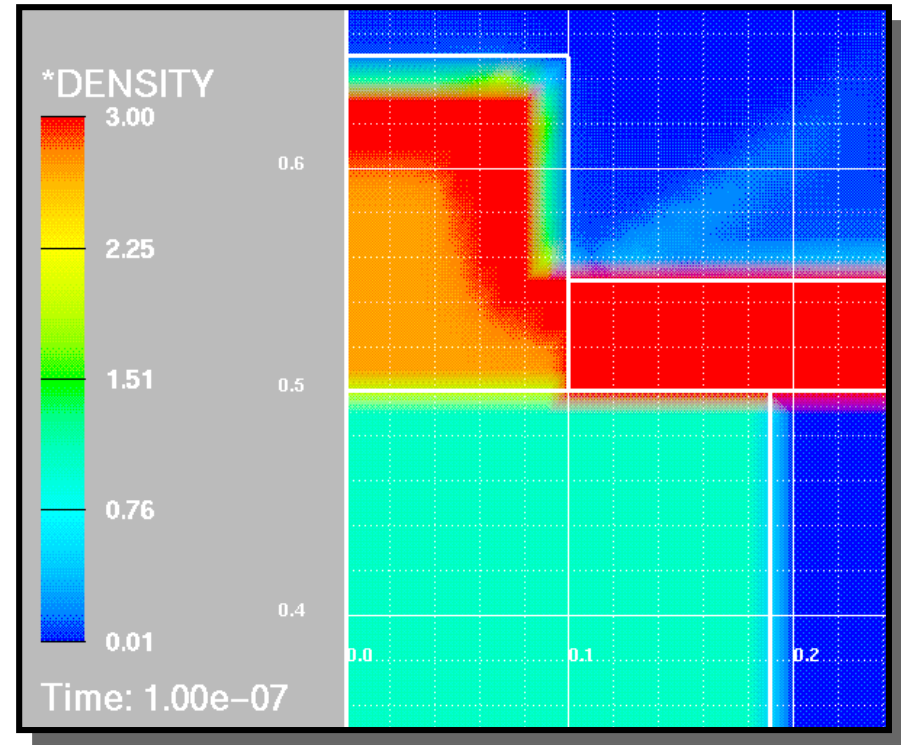
RadJet / NOVA:

rz-20-3



RadJet / Z:

rzj-30-3



Dimensions in cm





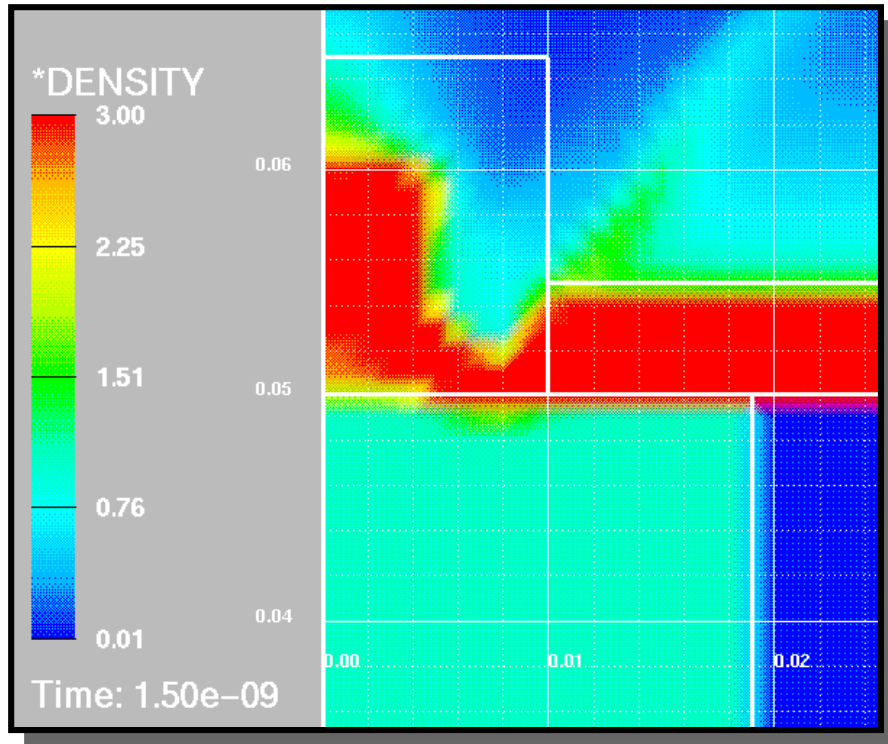
The RadJet problem at early times, at about the time of shock breakout . . .

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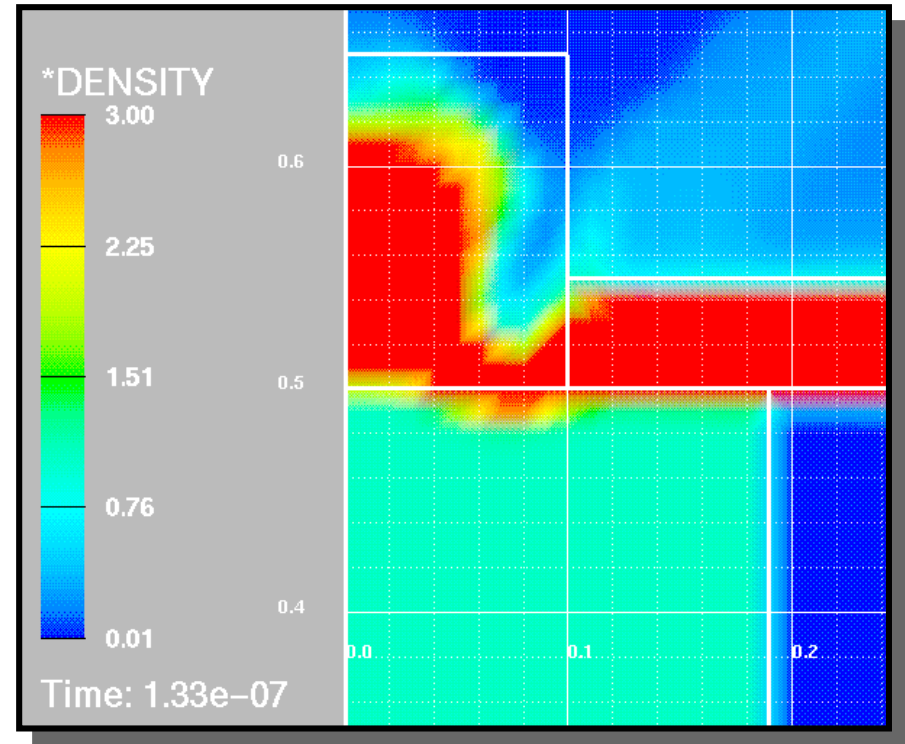
RadJet / NOVA:

ry-20-3



RadJet / Z:

ryz-30-3



Dimensions in cm





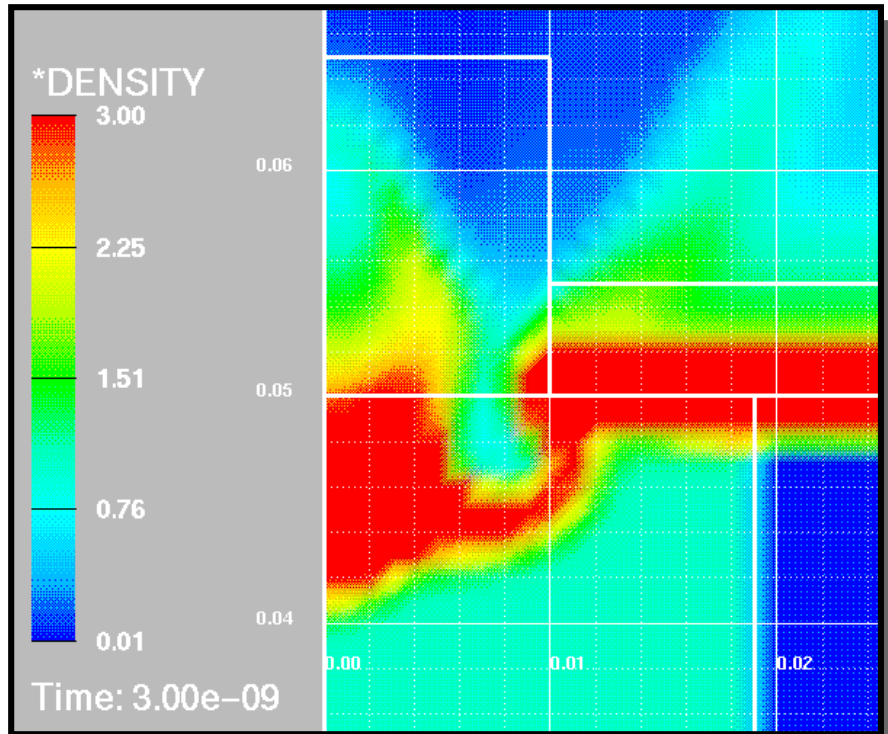
The RadJet problem at medium time, after the “jet” is relatively well formed . . .

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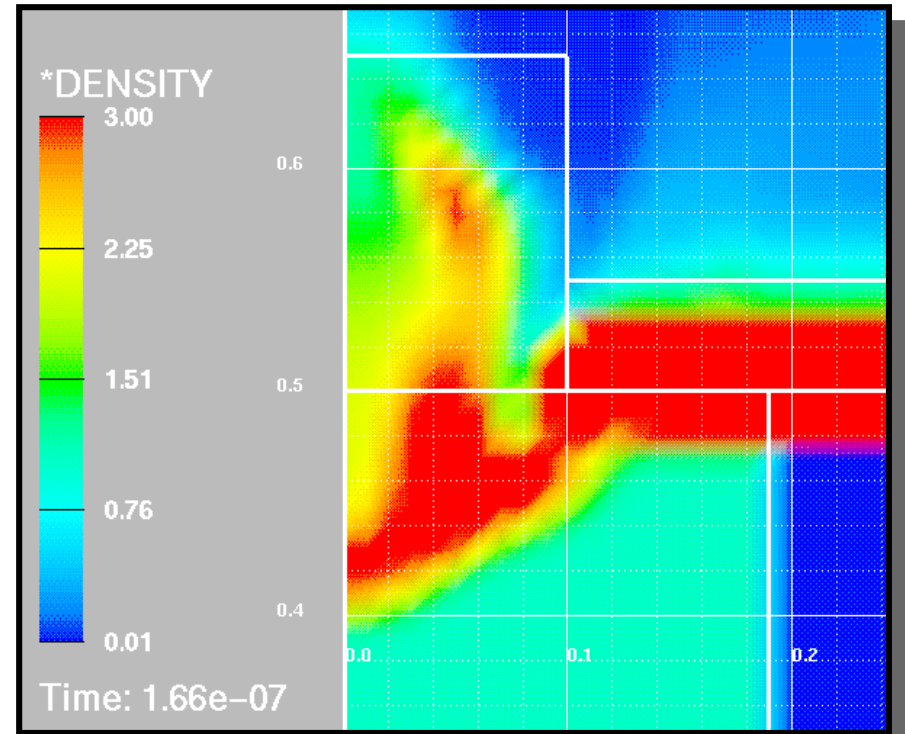
RadJet / NOVA:

ry-20-3



RadJet / Z:

ryz-30-3



Dimensions in cm





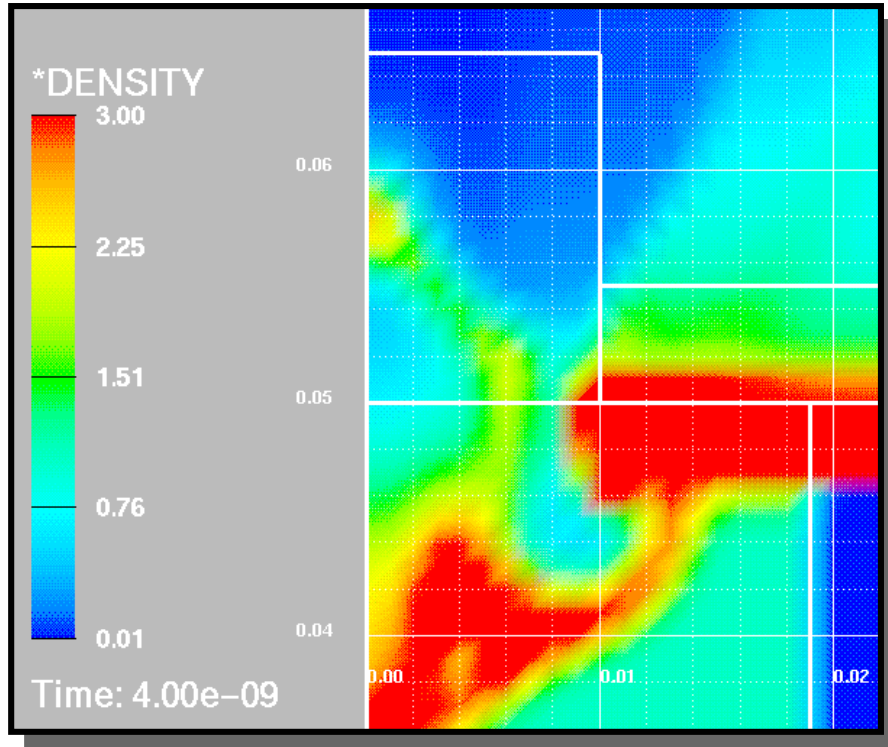
At later times the problems are still similar, and the jet is well into the polystyrene backing.

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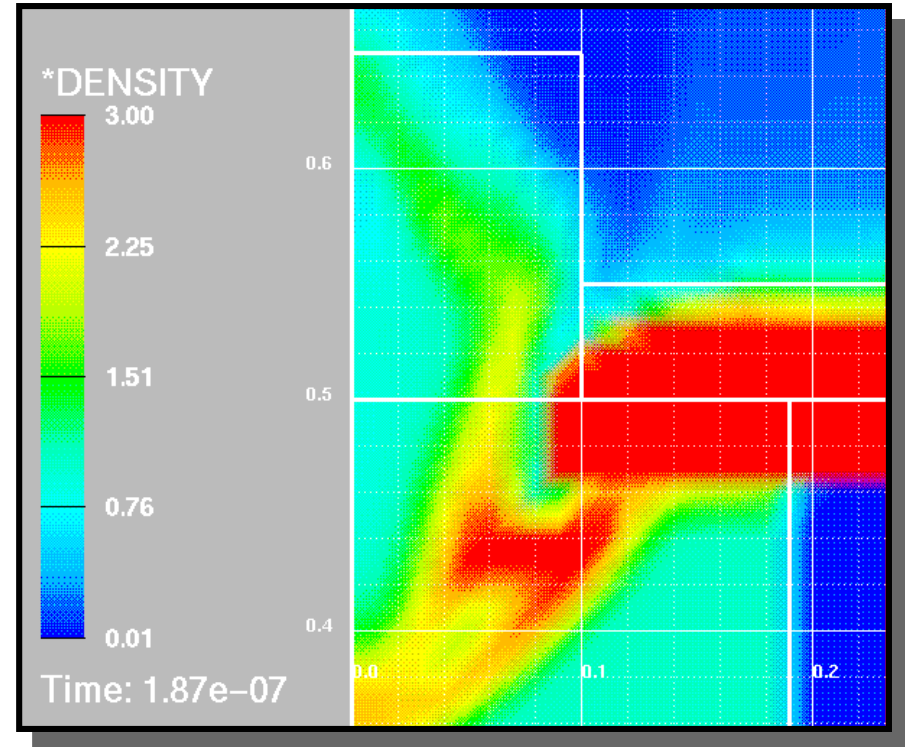
RadJet / NOVA:

ry-20-3



RadJet / Z:

ryz-30-6



Dimensions in cm





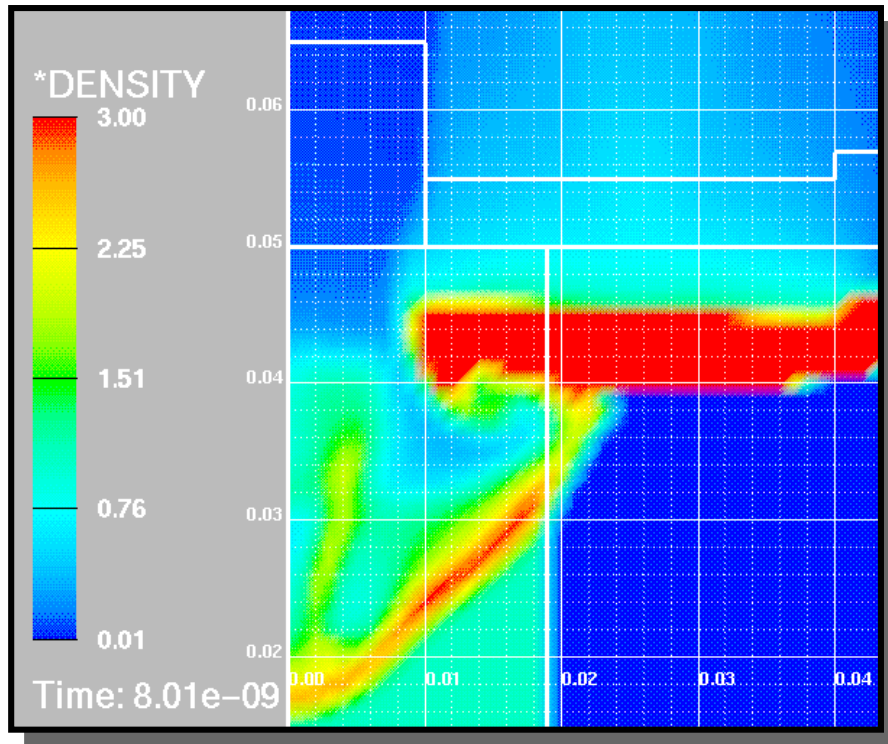
By late times, the differences due to lack of radiation “scaling” are considerably more evident.

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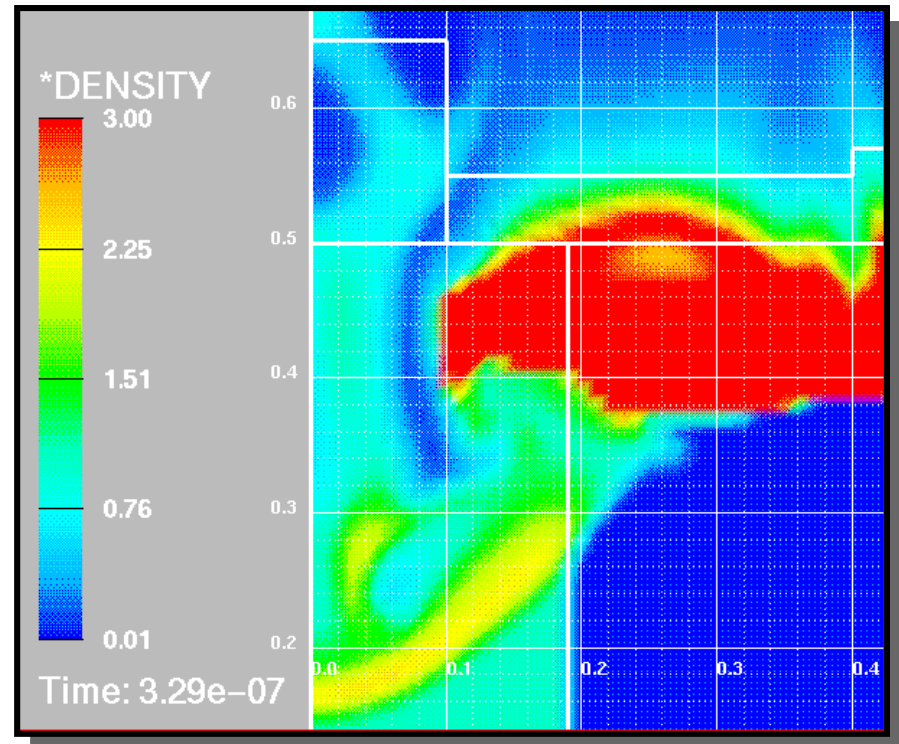
RadJet / NOVA:

ry-20-3



RadJet / Z:

ryz-30-6



Dimensions in cm





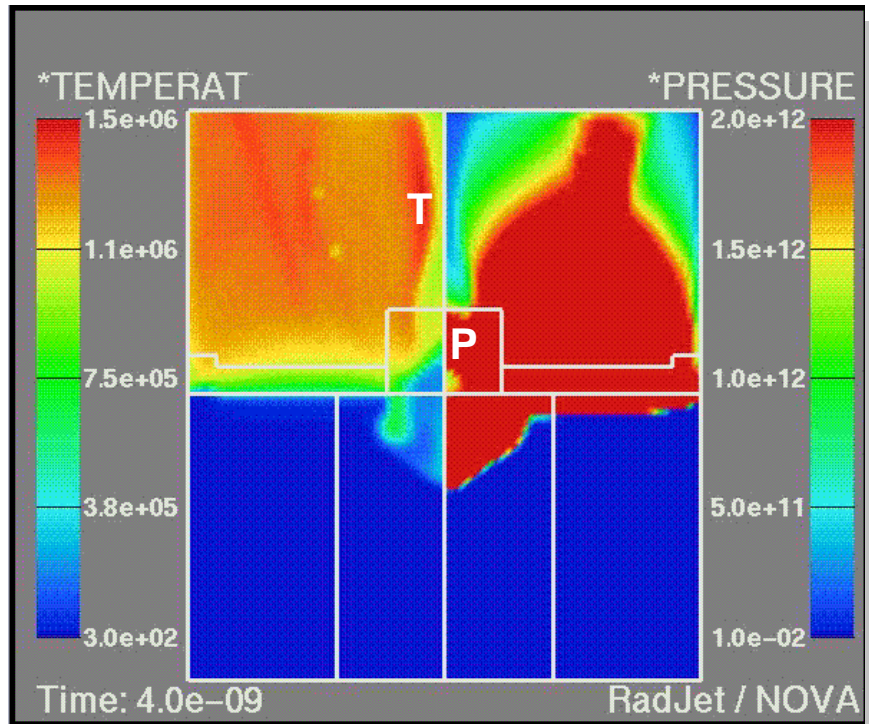
Temperatures and pressures show much larger differences than the densities.

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RadJet / NOVA:

ry-20-3

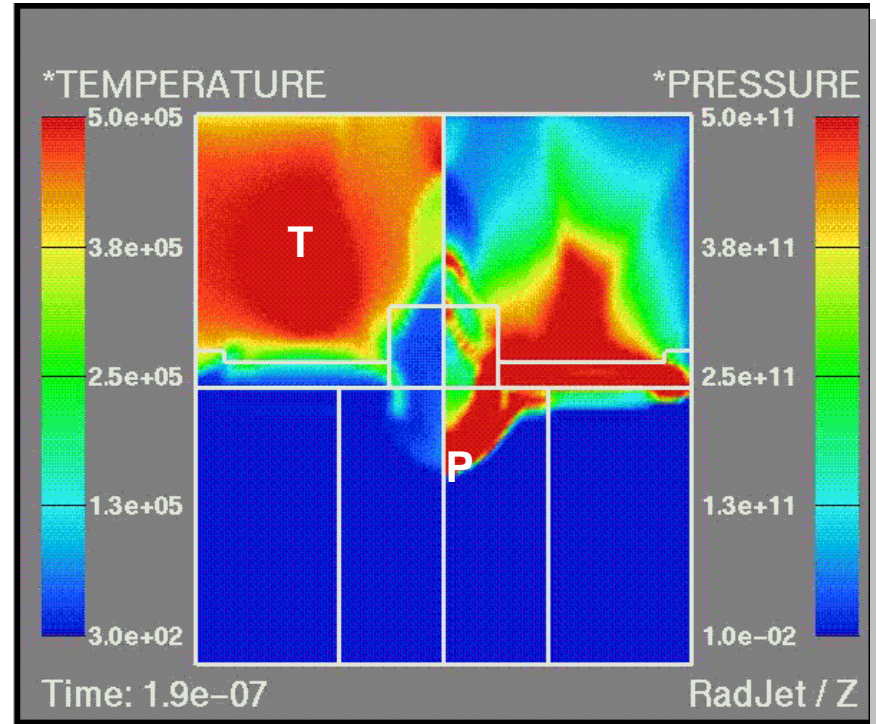


'T' -- $T_{\max} \cong 130 \text{ eV}$

'P' -- $P_{\max} \cong 30 \text{ Mb}$

RadJet / Z:

ryz-30-6



'T' -- $T_{\max} \cong 50 \text{ eV}$

'P' -- $P_{\max} \cong 3 \text{ Mb}$





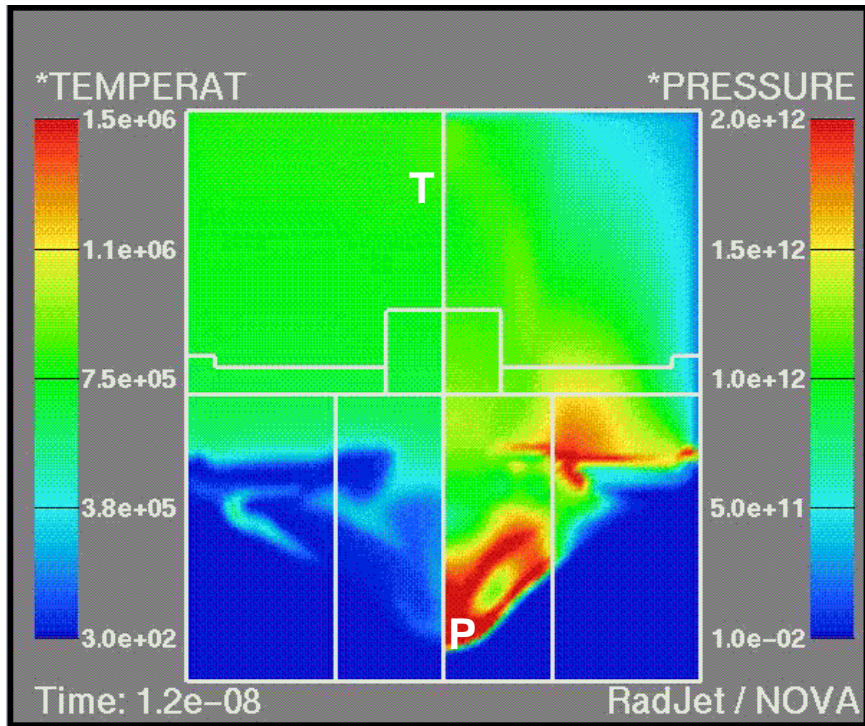
At much later times there are significant differences, but many qualitative features are similar.

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RadJet / NOVA:

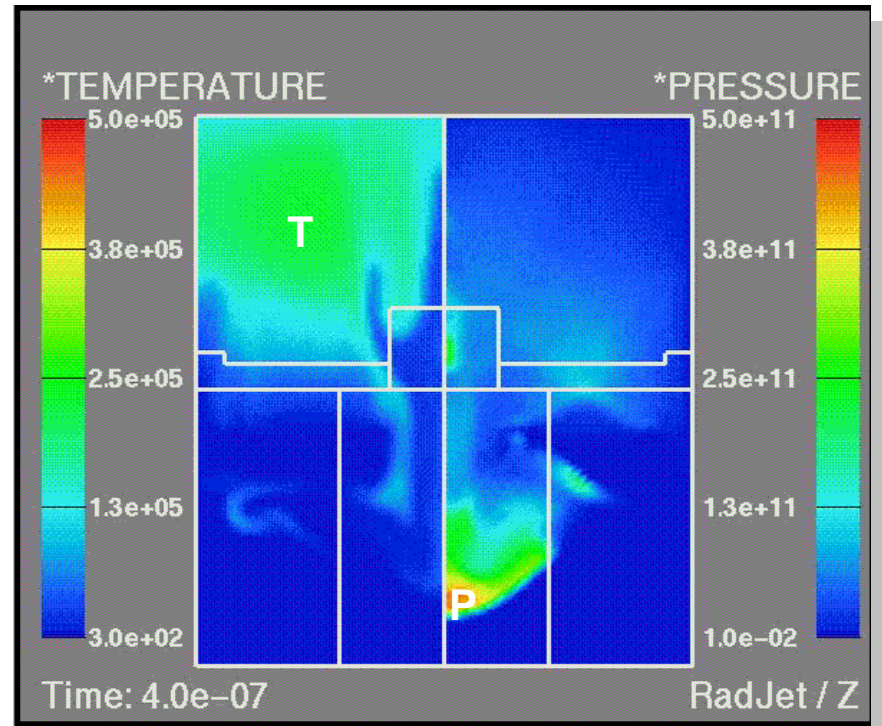
ry-20-3



'T' -- $T_{\max} \cong 70 \text{ eV}$
'P' -- $P_{\max} \cong 4 \text{ Mb}$

RadJet / Z:

ryz-30-6



'T' -- $T_{\max} \cong 20 \text{ eV}$
'P' -- $P_{\max} \cong 0.5 \text{ Mb}$





The Z-Beamlet Backlighter (ZBL) is scheduled to begin operation on Z in early 2001.

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- *Construction of the ZBL building began in March 1999.*
- *Construction of the ZBL building was completed in October 1999.*
- *The front end activation was completed in February 2000.*



March 1999



March 2000



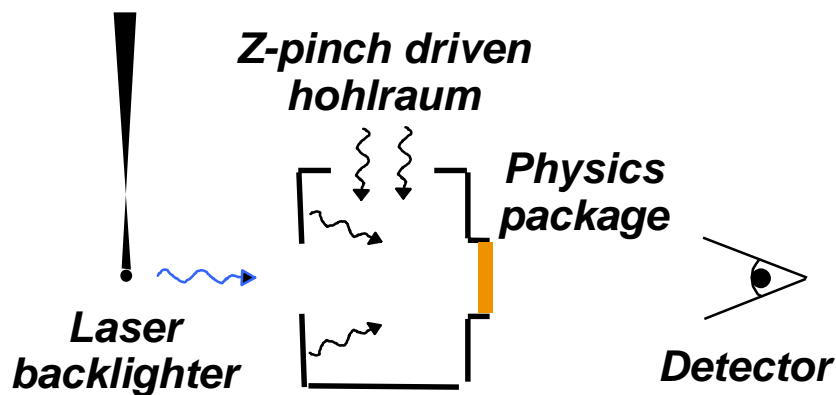
Sept 2000



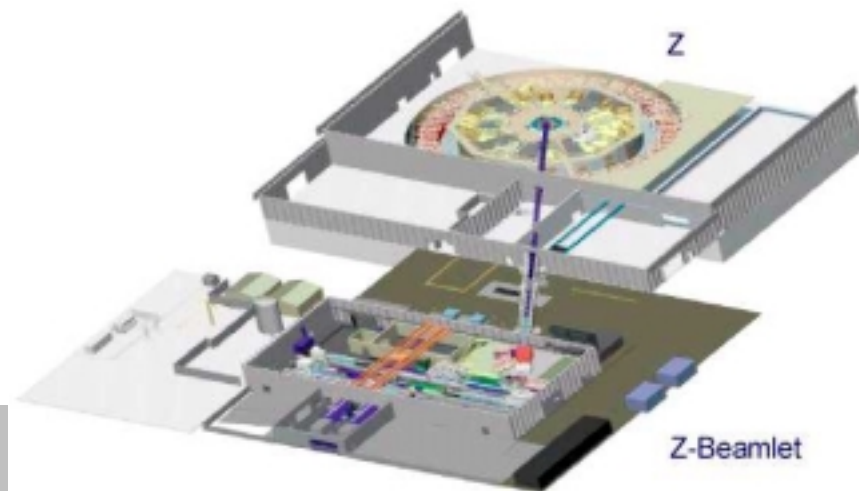


The ZBL will be an important new diagnostic tool for high energy density physics experiments on Z.

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2 TW laser backlighter on Z --



Measurements possible with a backlighter:

- Material T_e and n_e
- Magnetic Rayleigh-Taylor growth rate
- Absorption spectrum
- Capsule implosion symmetry
- Material interface motion
- Particle velocity and shock density
- Instability mix region

- Capabilities include both point projection and area backlighting.
- We will have spatial resolution of $25\ \mu\text{m}$ at 9 keV x-ray probe energy.



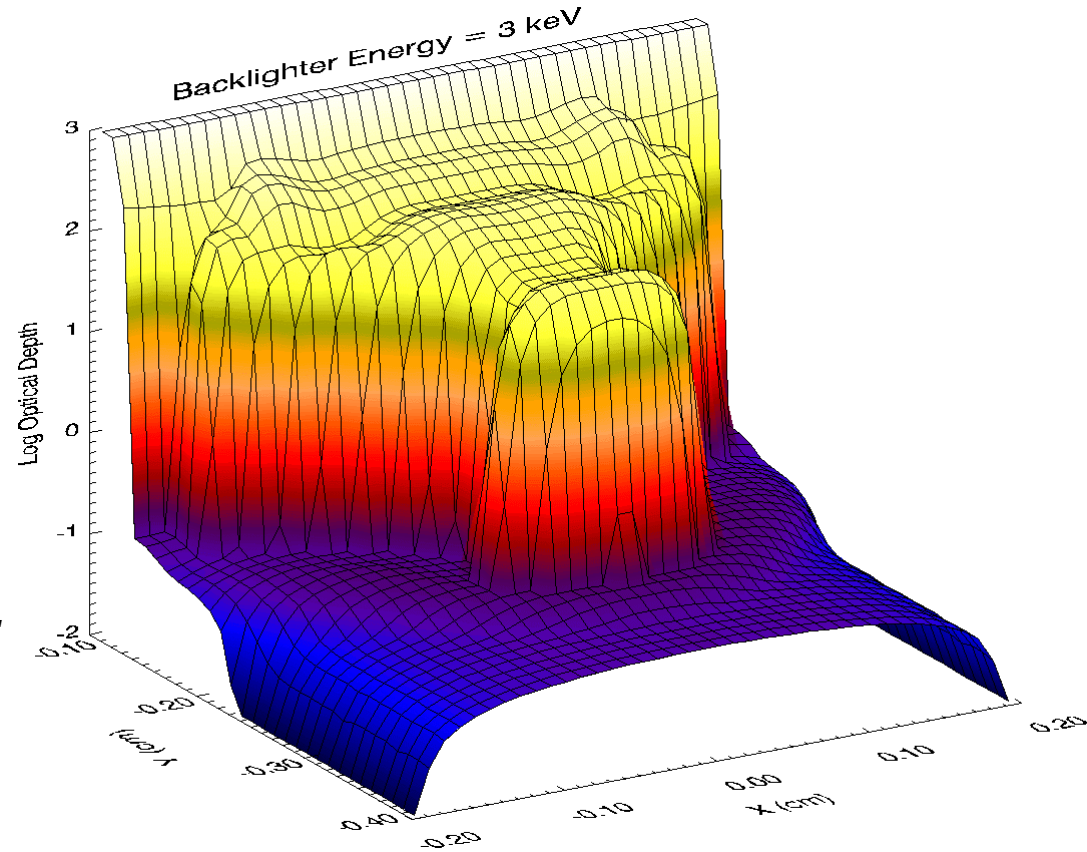


We are using SPECT3D to visualize the use of the Z-Beamlet Backlighter (ZBL) on these experiments.

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- *The amplitude of each cell represents the optical depth through the jet as a function of axial position (Y) and offset from the axis (X).*
- *Overall, ZBL performance depends on photon energy, conversion efficiency, and other issues.*
- *This plot is taken from the 100- μm resolution RadJet / Z calculation at a time of 330 ns.*
- *For this example the backlighter energy was chosen as 3 keV, but the jet is probably too thick to “see” through.*



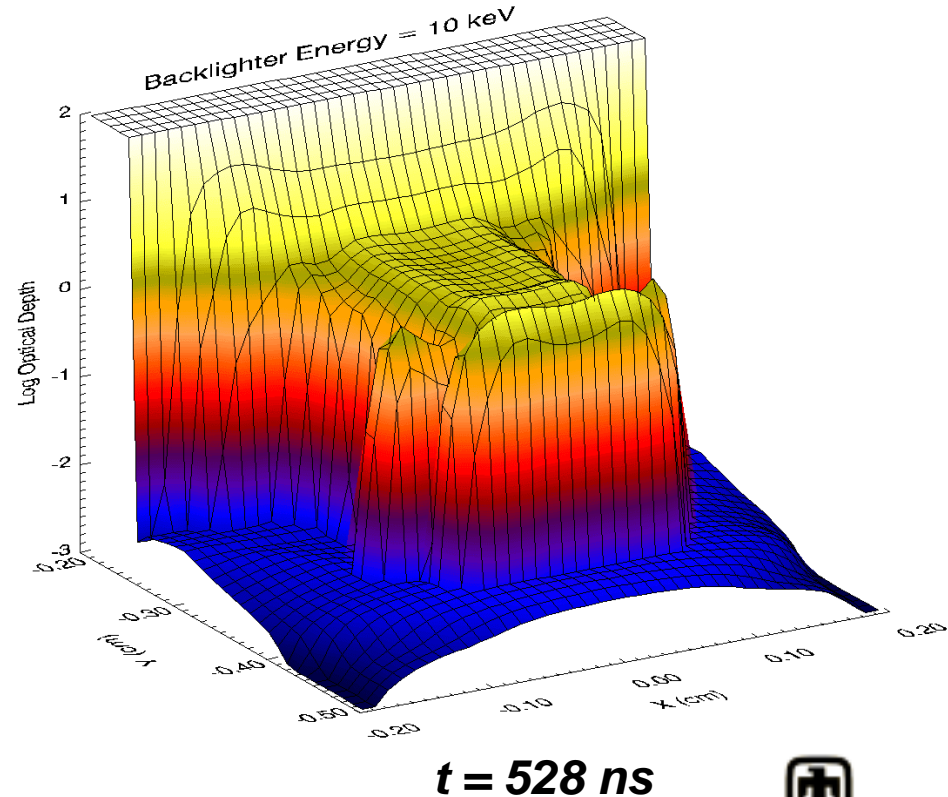
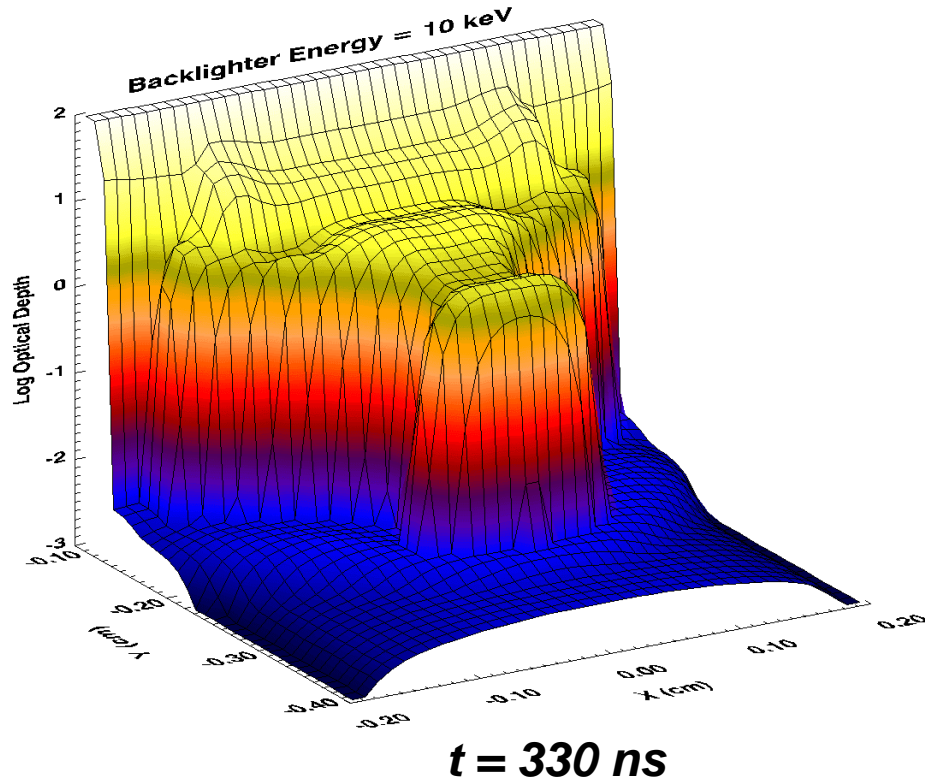


At late times and 10 keV, we get optical depths of order unity, which implies experiment is feasible.

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Backlighter photon energy:
 $h\nu = 10 \text{ keV}$



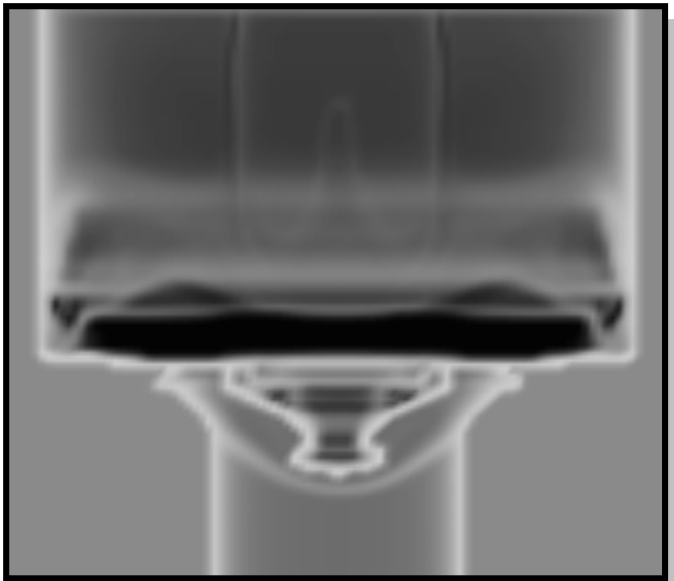


Simulations of detector output from the scaled-up Z runs show all major features.

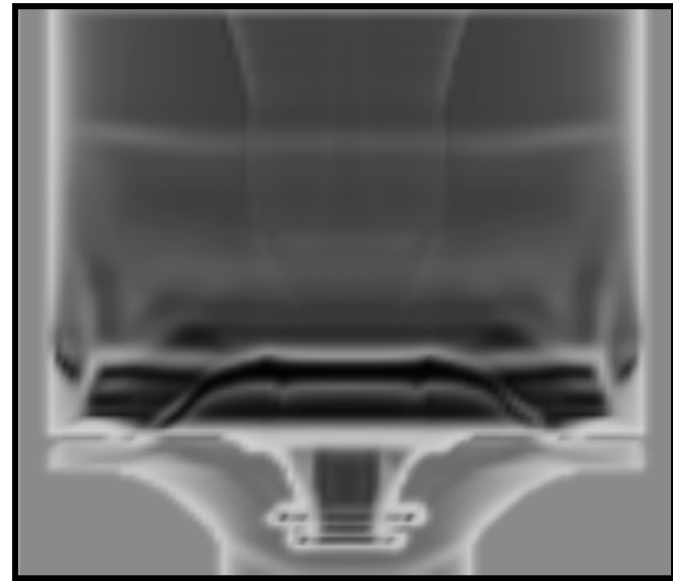
— Z-pinch and Target Theory Department —



- *These images were generated with $h\nu = 10$ keV.*
- *Features evident in the radiographs include the polystyrene backing block, the shock wave in the polystyrene, and details of the aluminum jet in the plastic.*
- *Details of the blowoff moving back into the hohlraum are also evident, but would not be recorded in the experimental radiograph.*



$t = 330$ ns



$t = 528$ ns





We have studied the generation and evolution of radiation-driven jets on both NOVA and Z.

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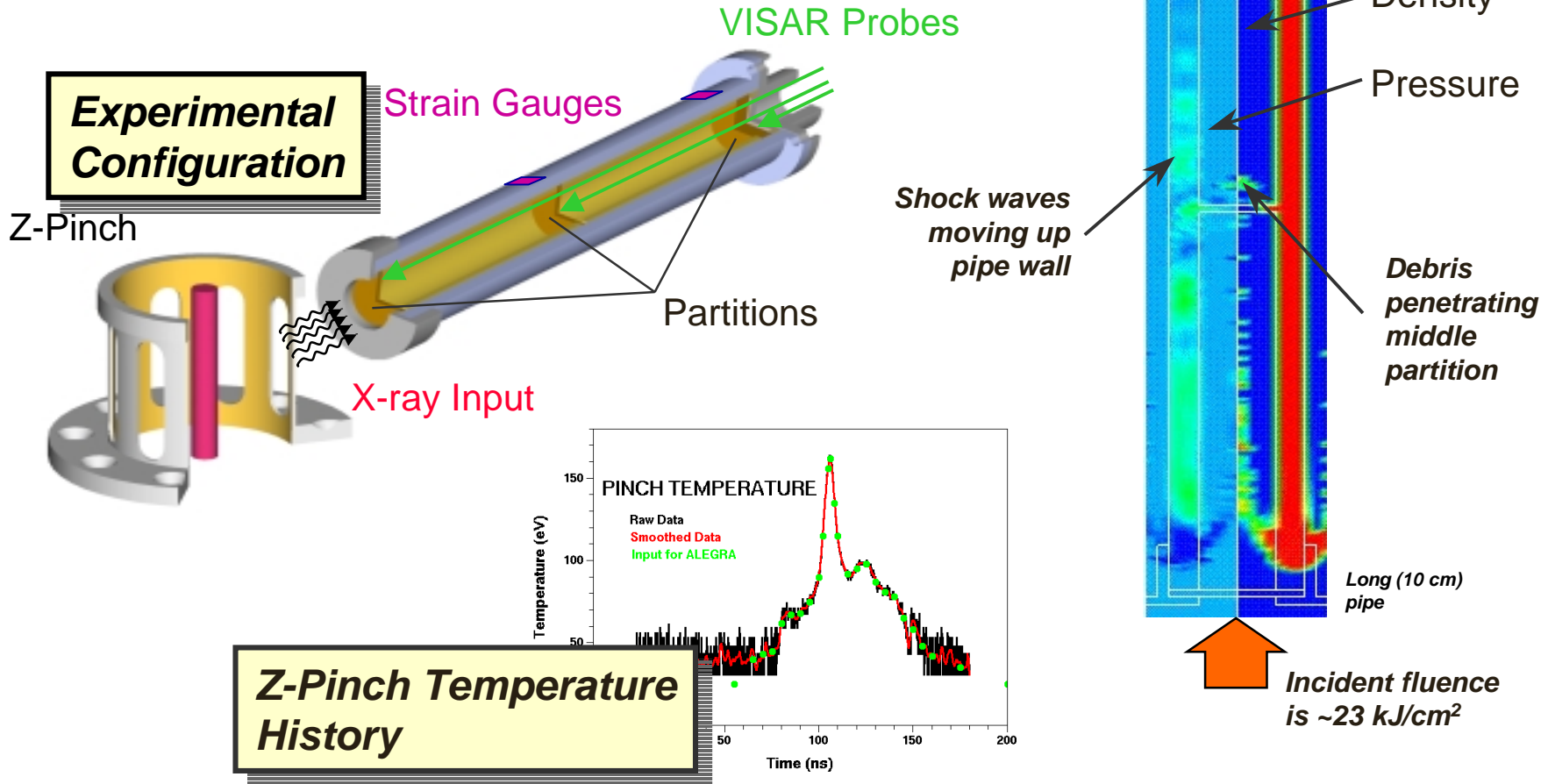
- The NOVA experiments, in conjunction with the other calculations, have provided validation for the ALEGRA modeling and analyses.
- In comparison with the results from NOVA, physical scaling-up of the configuration and using the Z-pinch machine produces similar, although not identical, phenomenology.
- Using the ZBL backlighter for diagnostic measurements appears to be feasible for the scaled-up configuration.
 - > At late times and for high photon energies, optical depths are of order unity.
- Next steps and other possibilities:
 - > Use finer zoning for ALEGRA calculations;
 - > Use more realistic and representative ZBL spectra;
 - > Modify Z source to obtain different conditions (e.g., higher temperatures via dynamic hohlraum, multiple and/or colliding jets);
 - > Examine different configurations of interest, or other degrees of physical scale-up.



Debris Effects in Partitioned Pipes

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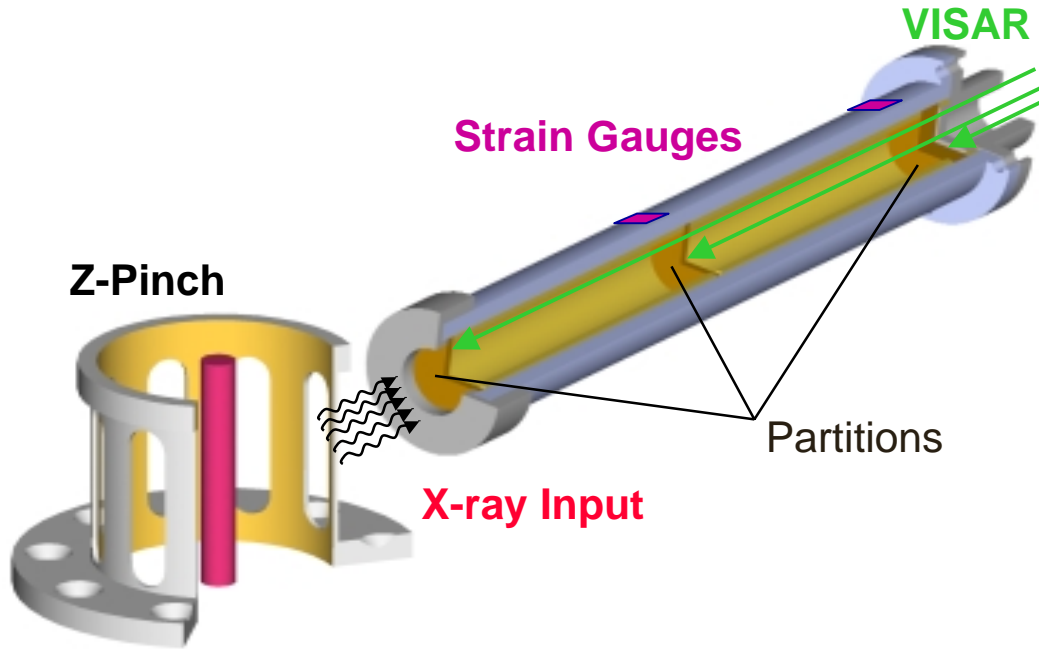
**ALEGRA Calculation
at $t = 15 \mu s$**





This experiment is being used to study pipe and debris phenomenology using Z and ALEGRA.

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- *Well-characterized radiation environments*
- *Two configurations -- 10 cm and 5 cm long x ~1 cm*
- *Multiple shock diagnostics*
- *Real-time debris generation and barrier failure*

Strategic Intent:

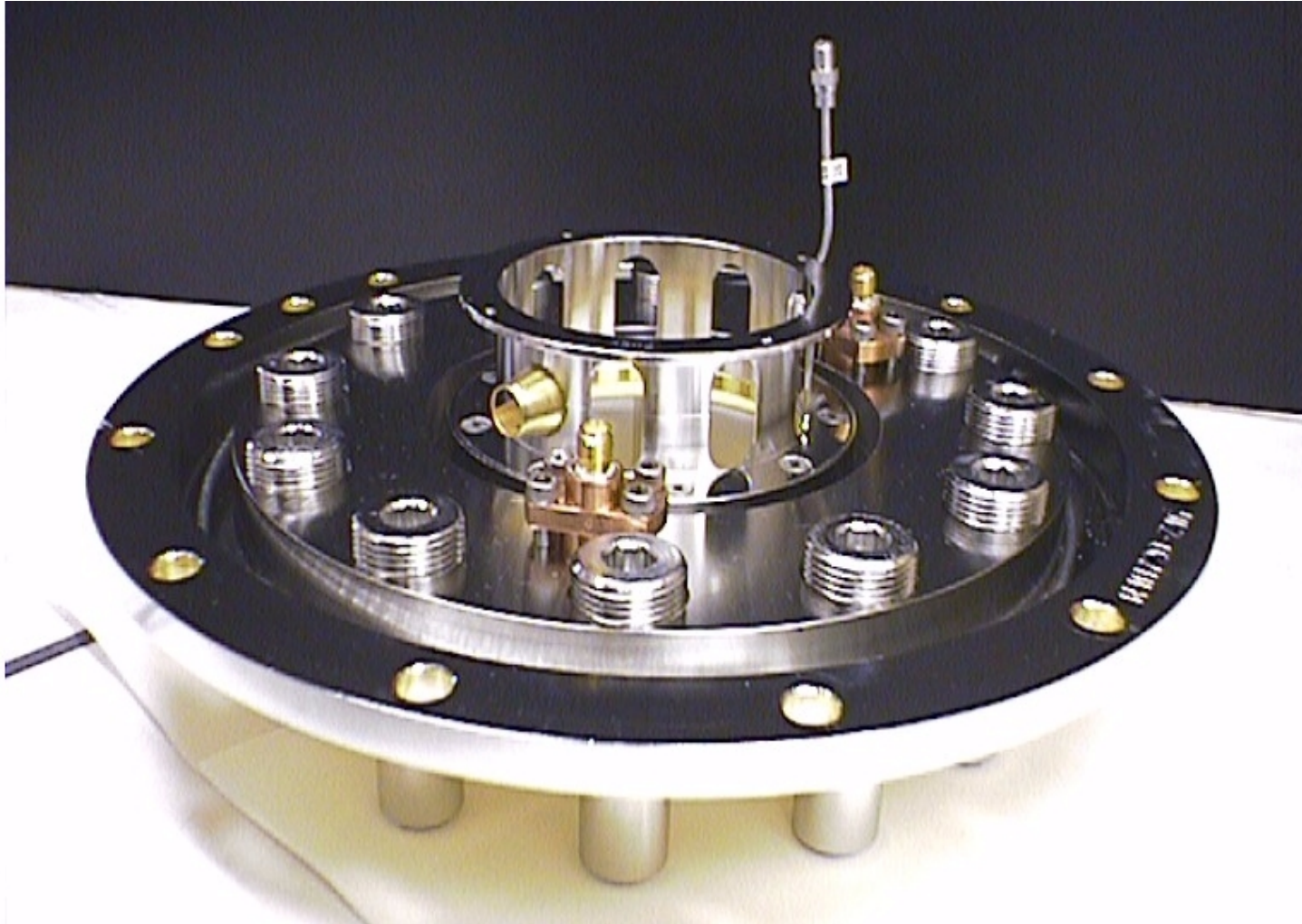
- Show new capability for Z and Z-pinch in weapon-related applications;
- Demonstrate capabilities and expertise of interest to other agencies;
- Develop new techniques and diagnostics for future applications, and new areas for ASCI code validation with ALEGRA.





This slotted-can hohlraum configuration is being used for the partitioned-pipe experiments.

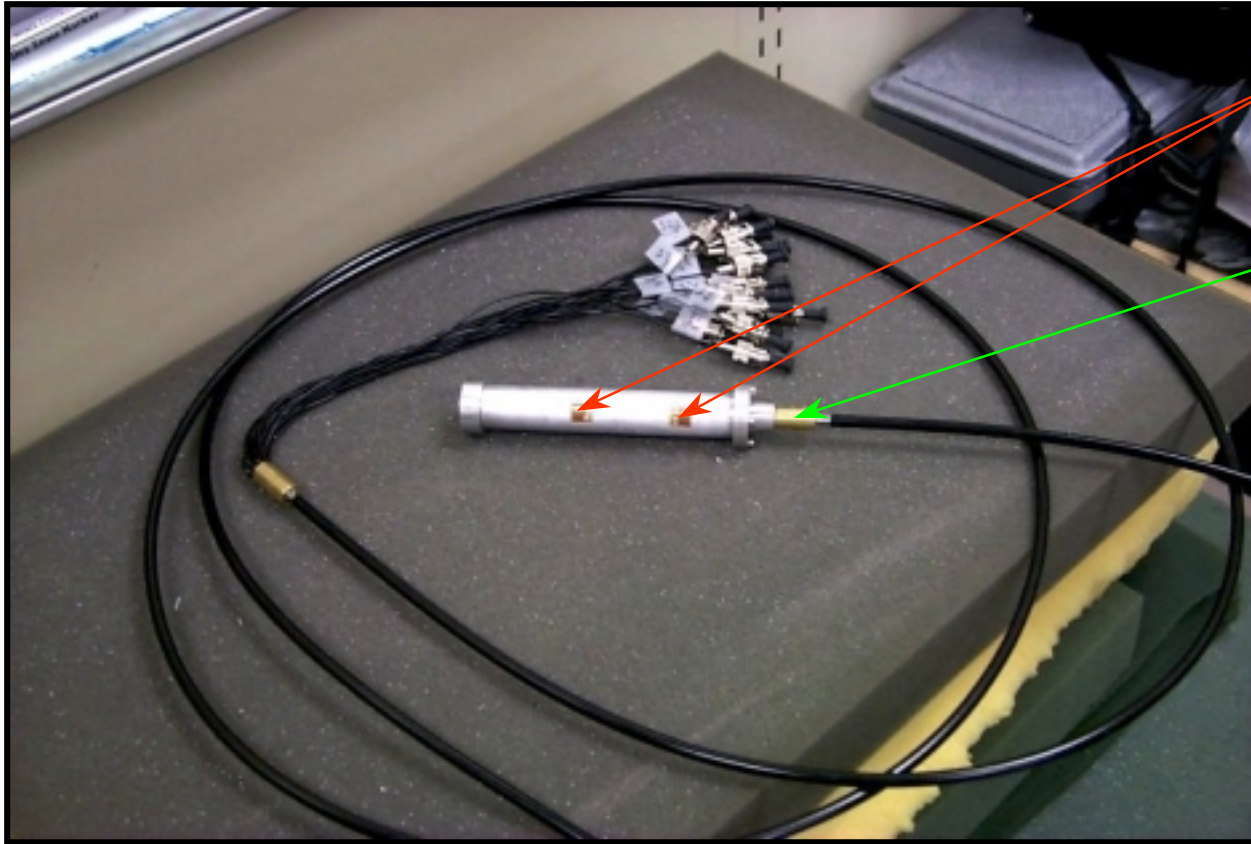
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The actual long-pipe sample shows the strain gauges and the attached VISAR probes.

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Strain gauges

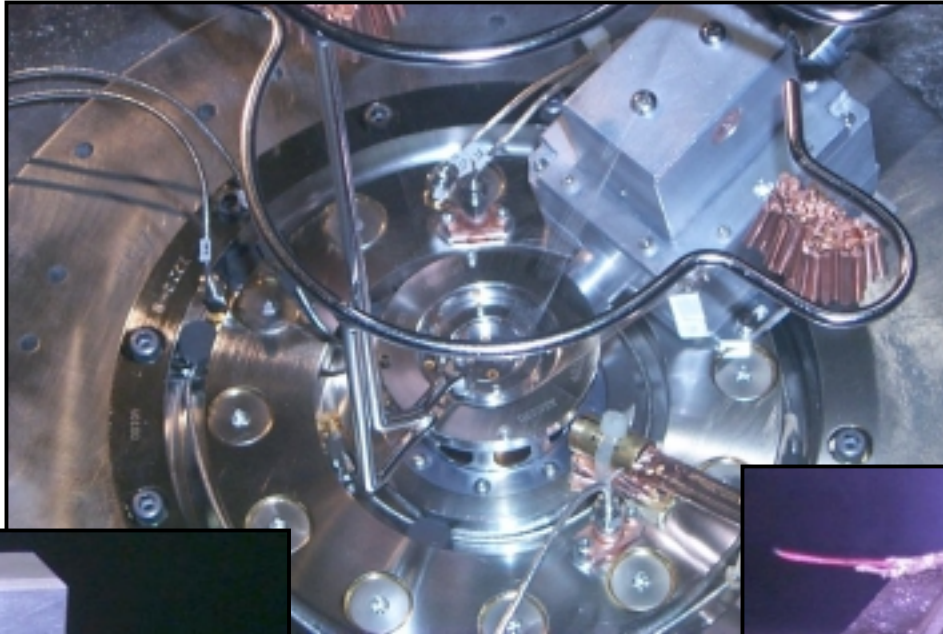
*Connections for
VISAR probes*



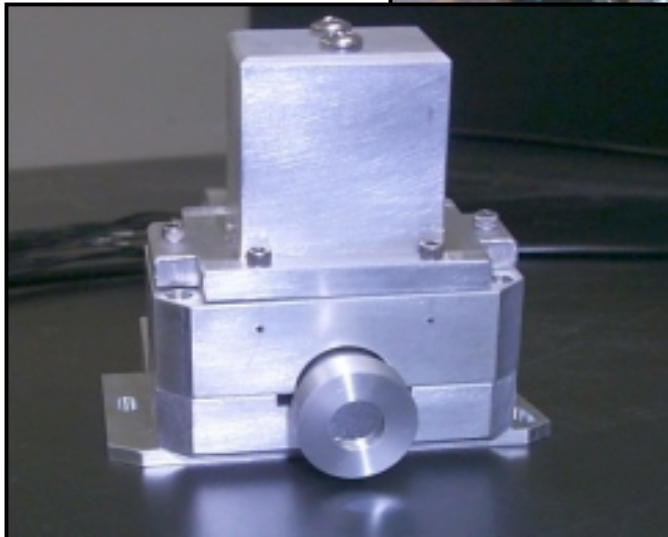


The experiment was successful, and VISAR data was recorded for all three partitions.

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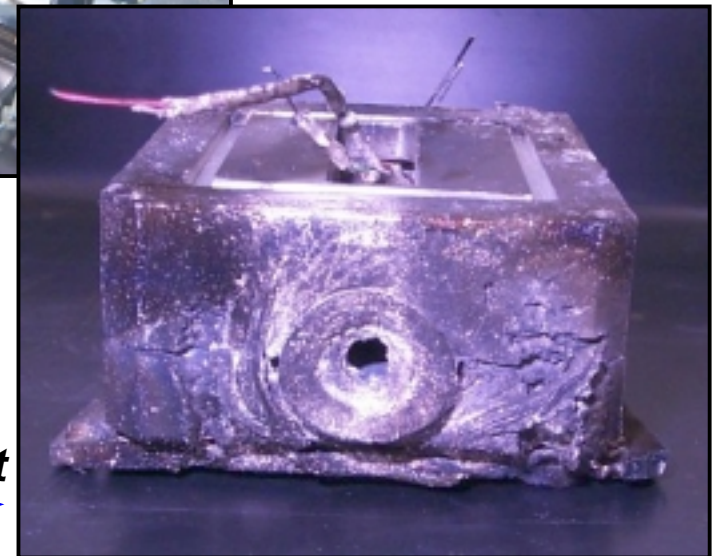
Model target assembly mounted in Z



Tunnel model in experimental fixture

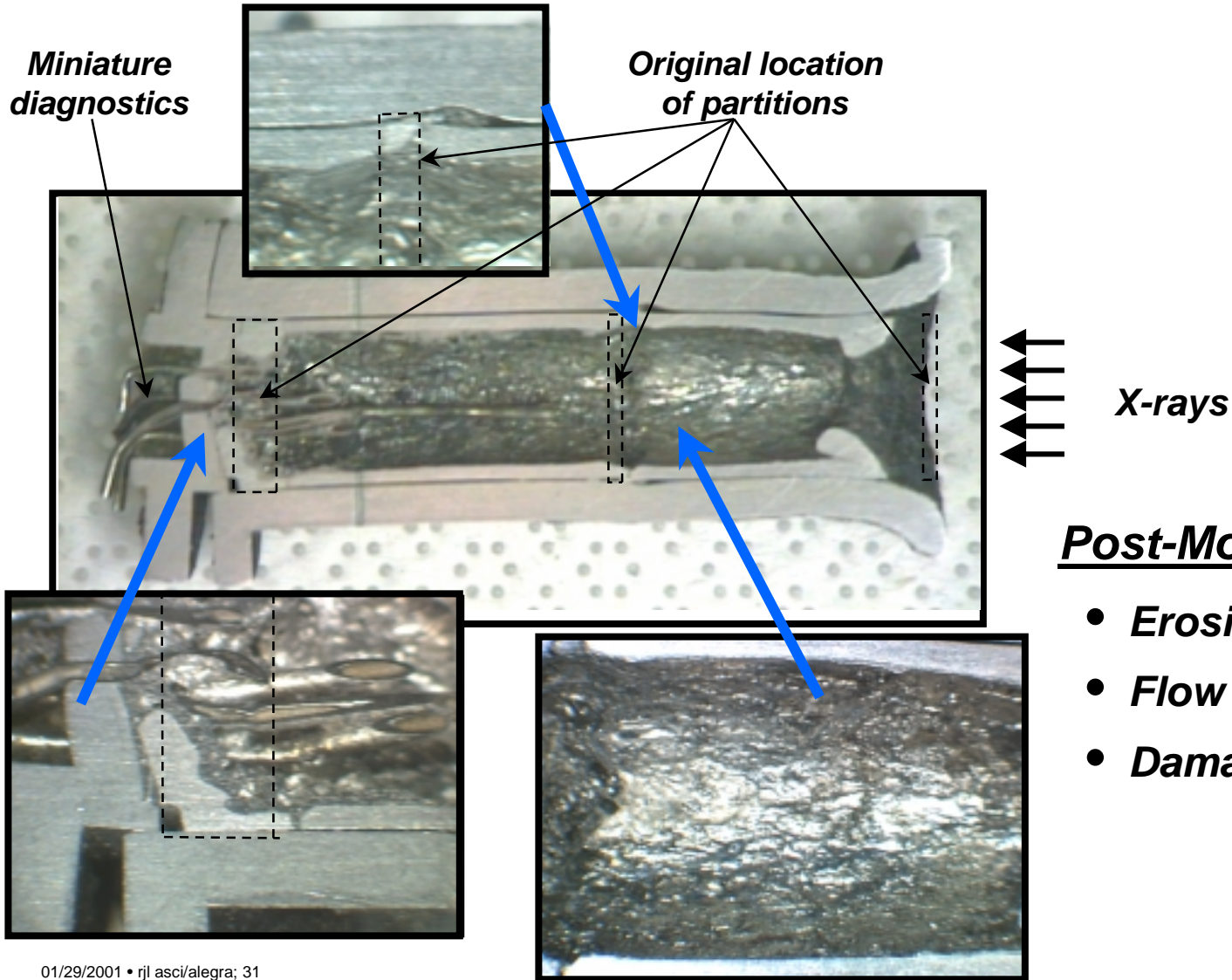
Post-shot

Pre-shot



Recovered short pipe provides evidence of response and confirmation of code simulations.

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Post-Mortem Analysis

- *Erosion mechanisms*
- *Flow phenomenology*
- *Damage mechanisms*

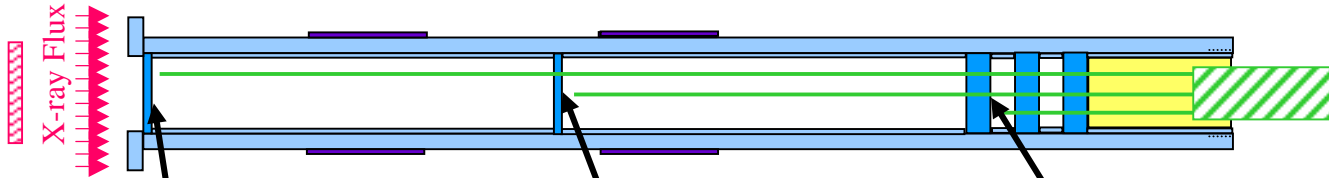


Velocity measurements functioned well, and data were obtained for all partitions.

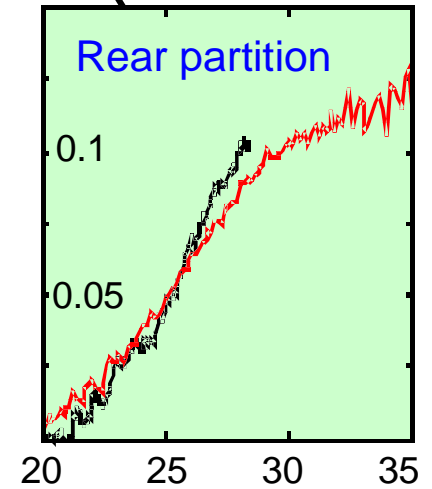
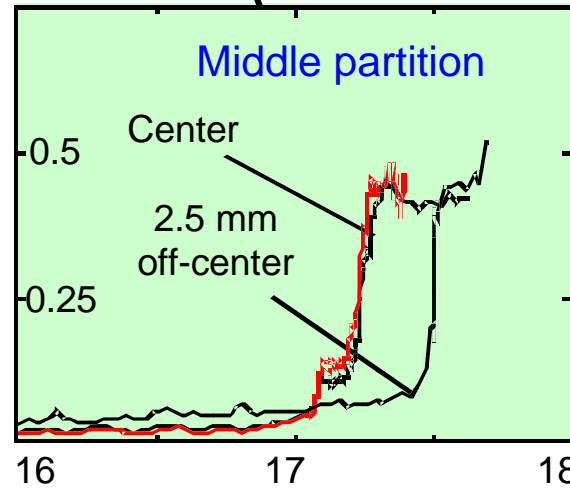
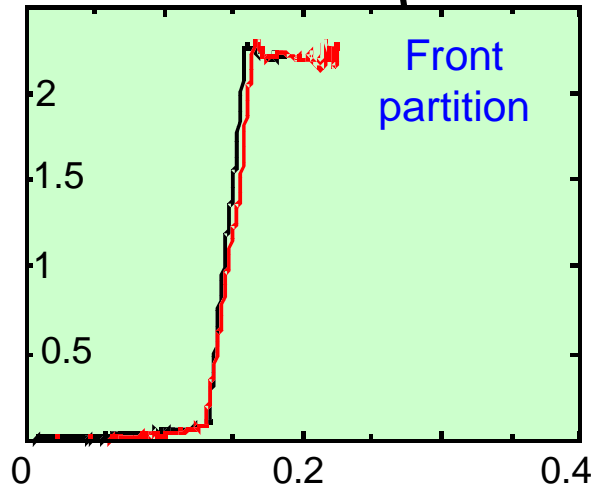
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Long Pipe



Partition Velocity (km/s)



Time (μs after pinch)

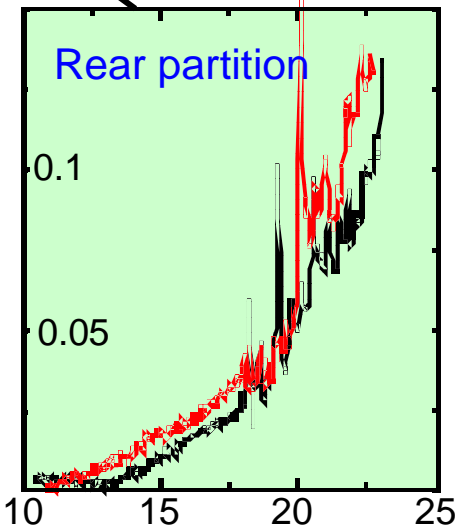
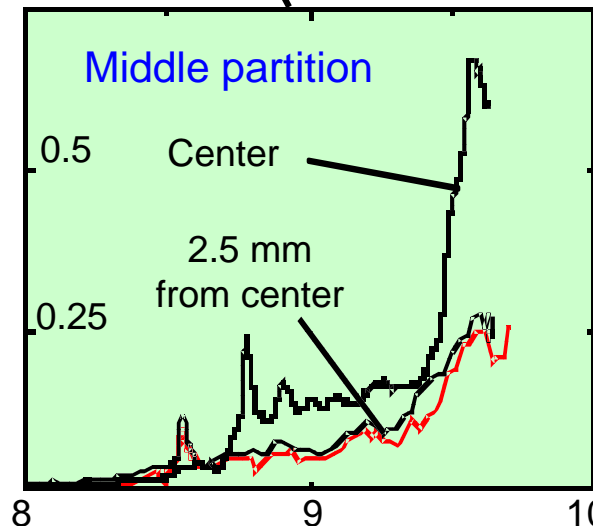
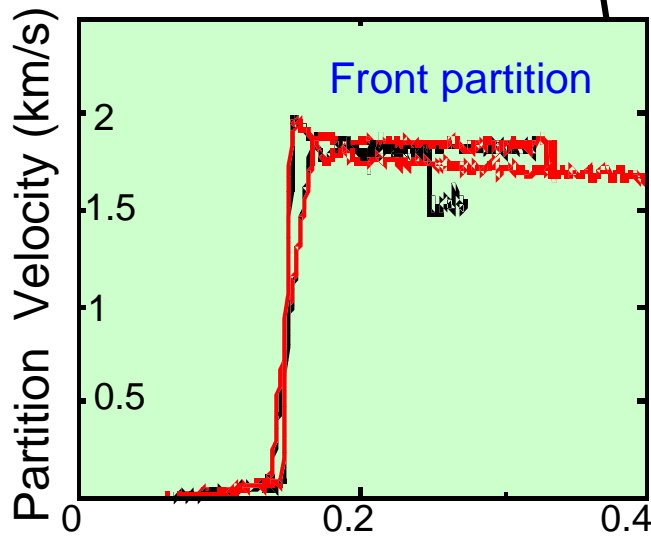
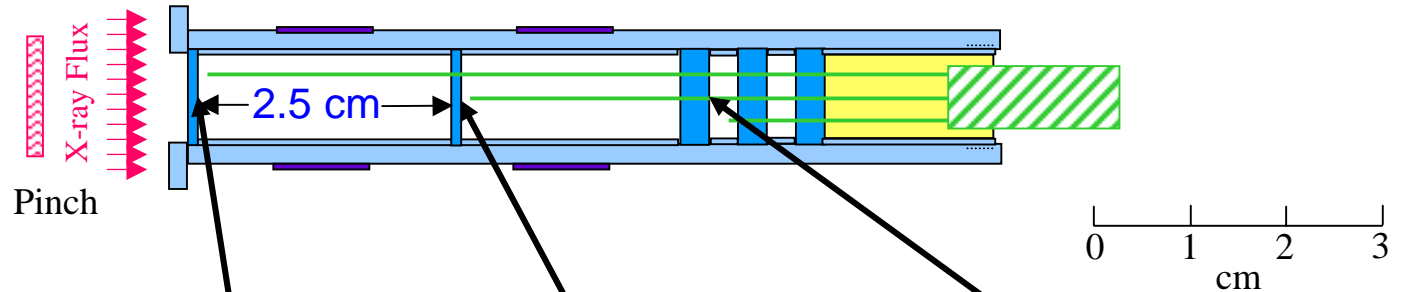


However, the data do show significant scatter, indicating the variable nature of the phenomena.

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Short Pipe



Time (μ s after pinch)

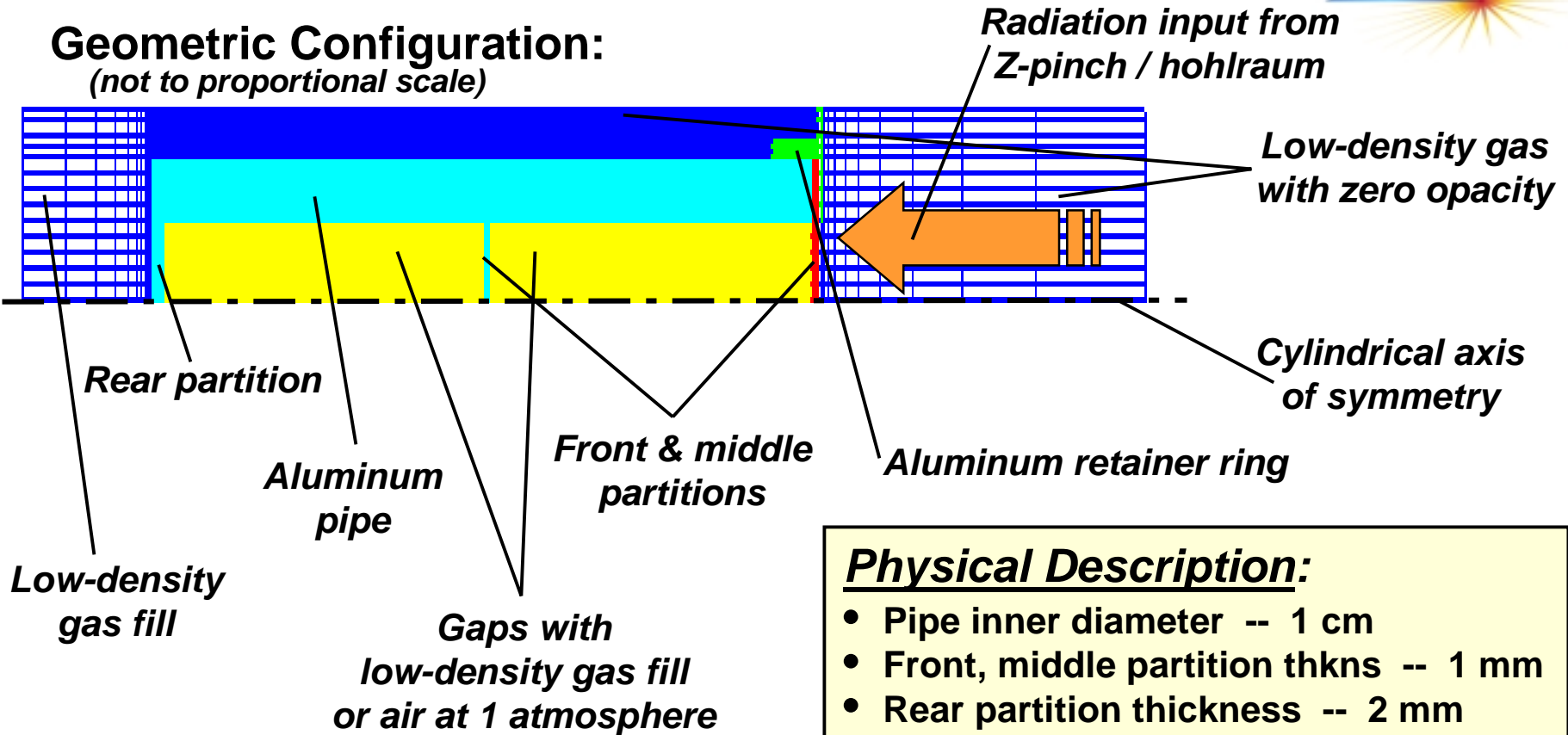


We are using ALEGRA for 2-D Eulerian simulations (post-shot) of the partitioned-pipe response.

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Geometric Configuration: (not to proportional scale)



“Long” Pipe

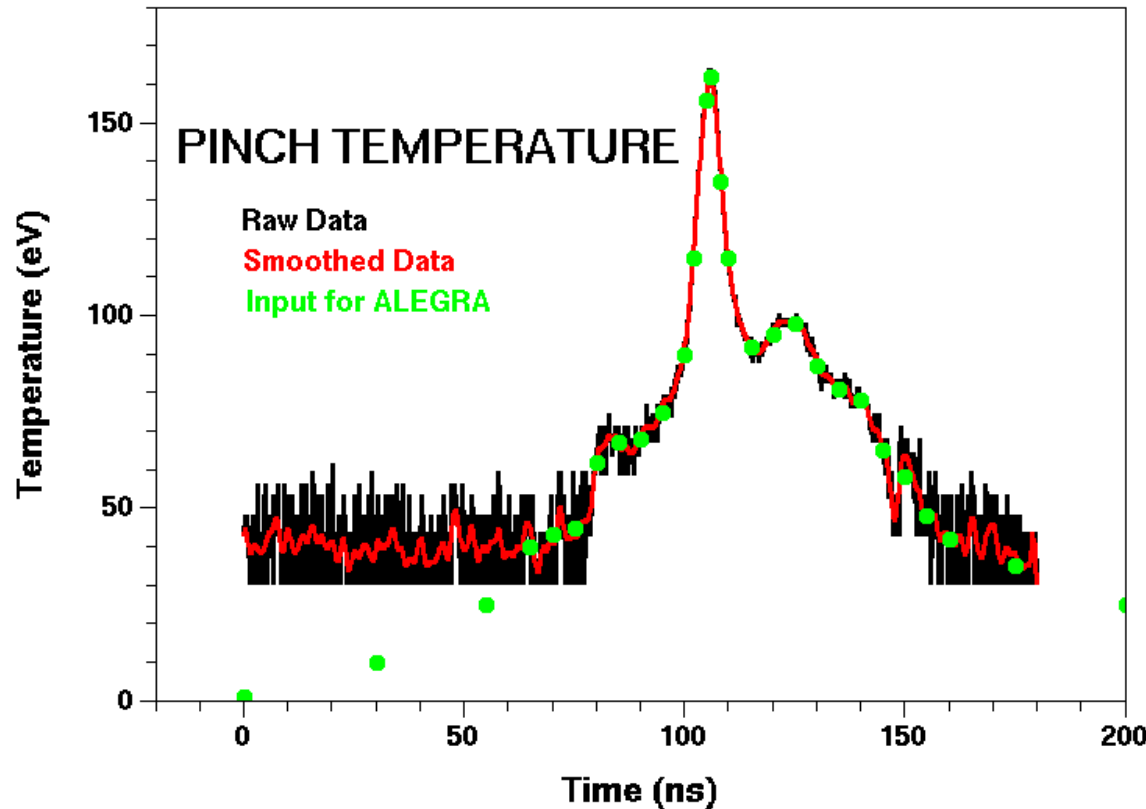
Physical Description:

- Pipe inner diameter -- 1 cm
- Front, middle partition thkns -- 1 mm
- Rear partition thickness -- 2 mm
- Wall thickness -- ~ 4 mm
- Gap length -- 50 mm
- Mesh -- ~ 7,000 elements



The measured output from the Z pinch is being used as input to the post-shot calculations.

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- *The pre-shot estimate was a 40-ns wide triangular pulse peaking at ~150 eV.*
- *The measured output has a peak temperature of ~162 eV, with the indicated time dependence.*
- *Geometric factors restrict the on-target fluence to about 3.3 percent of an equivalent hohlraum at the pinch temperature.*
- *Radiation transport used single-group diffusion; probably adequate for this problem.*

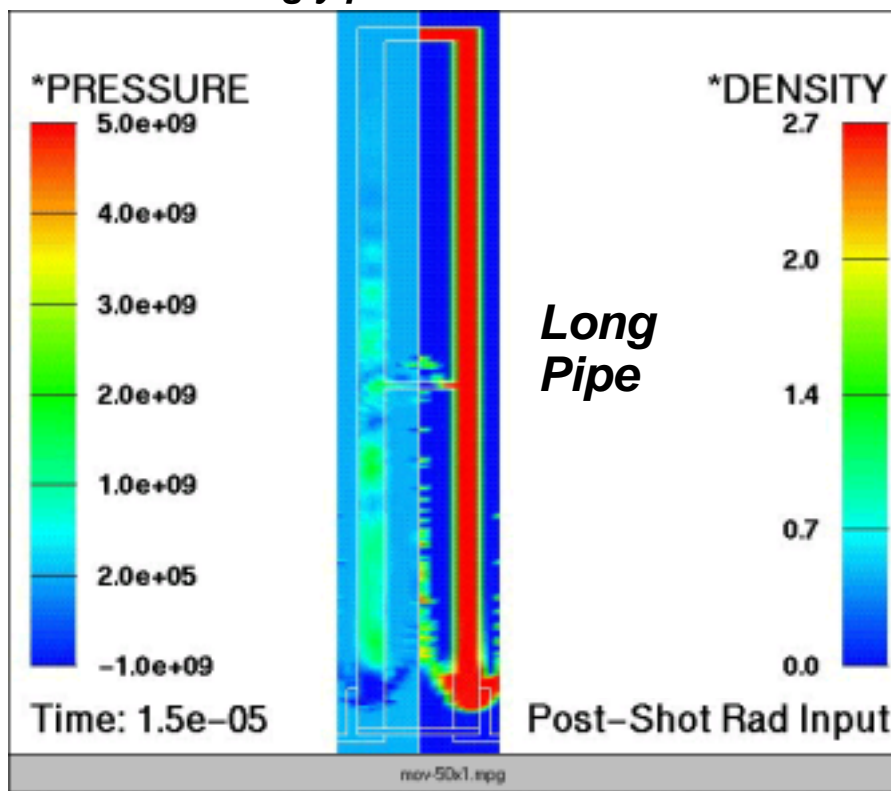


Post-shot calculations were performed with measured radiation input and E-P mat'l response.

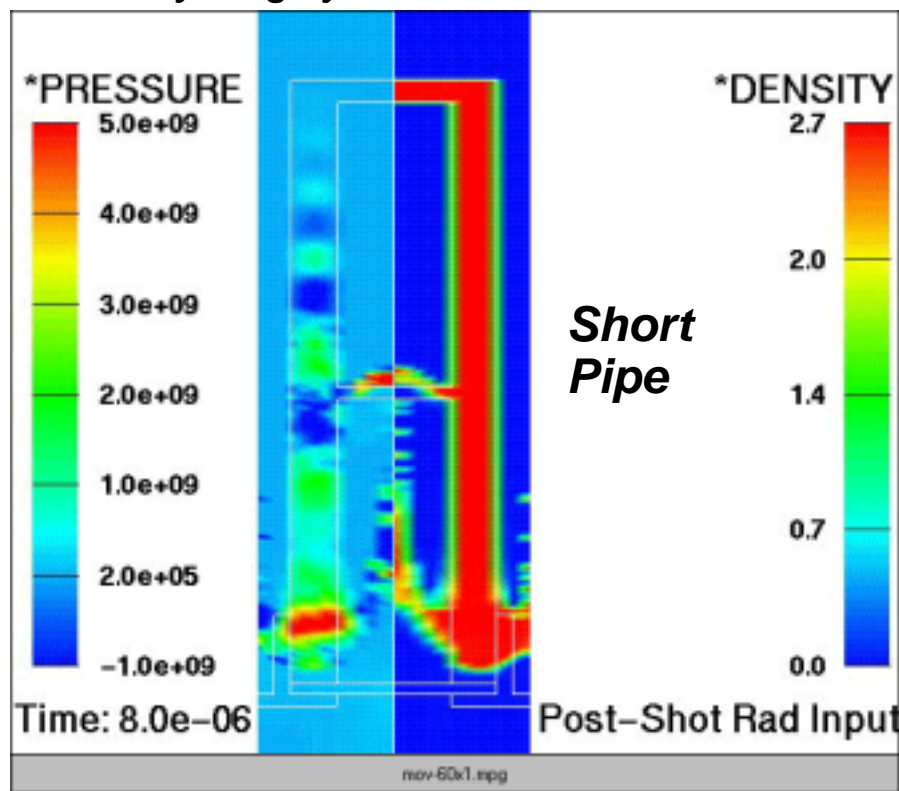
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The long-pipe calculation shows penetration of the middle barrier at $\sim 13 \mu\text{s}$ and penetration of the rear barrier at $\sim 36 \mu\text{s}$. In both problems the debris is strongly peaked toward the axis.

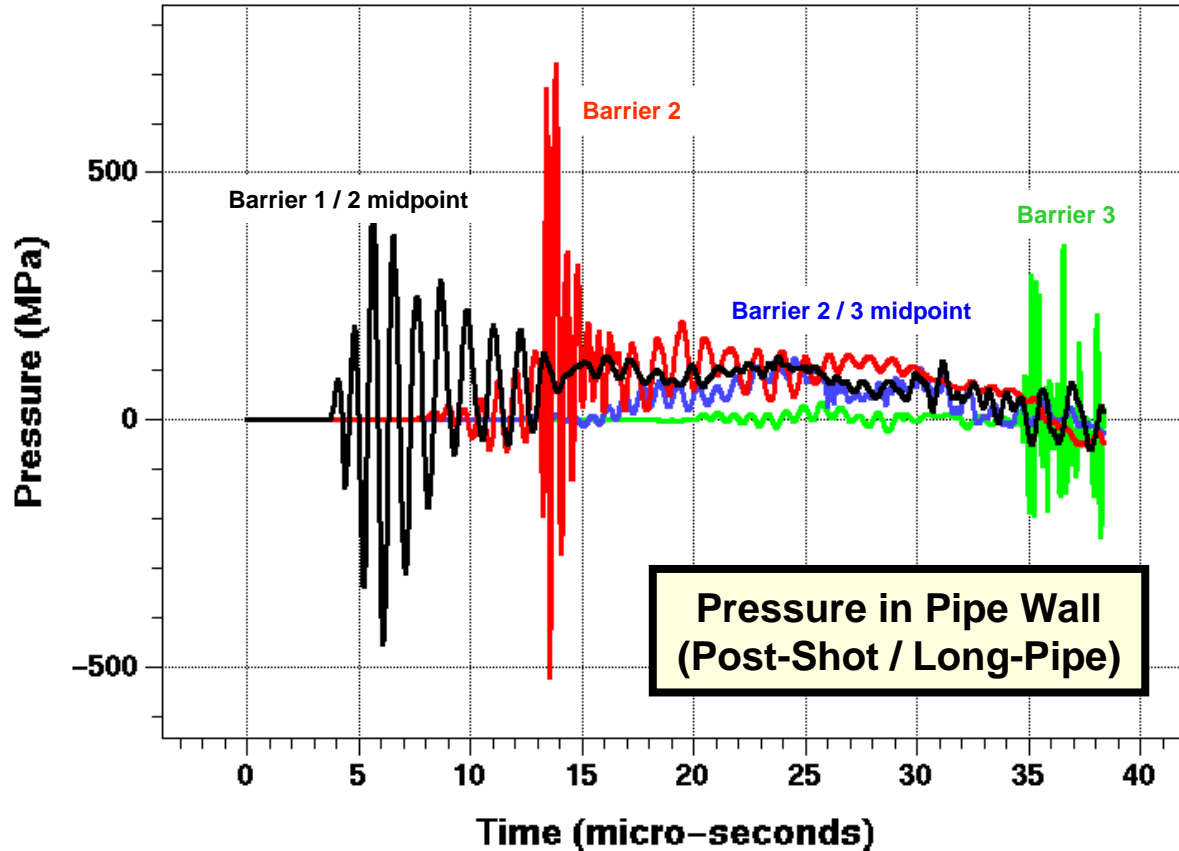


The short-pipe calculation shows penetration of the middle barrier at $\sim 7 \mu\text{s}$, and penetration of the rear barrier at $\sim 20 \mu\text{s}$. In both cases the stress waves in the tunnel walls precede the debris by roughly a factor of two in time.



ALEGRA calculations show numerical simulations for pressure histories along pipe wall.

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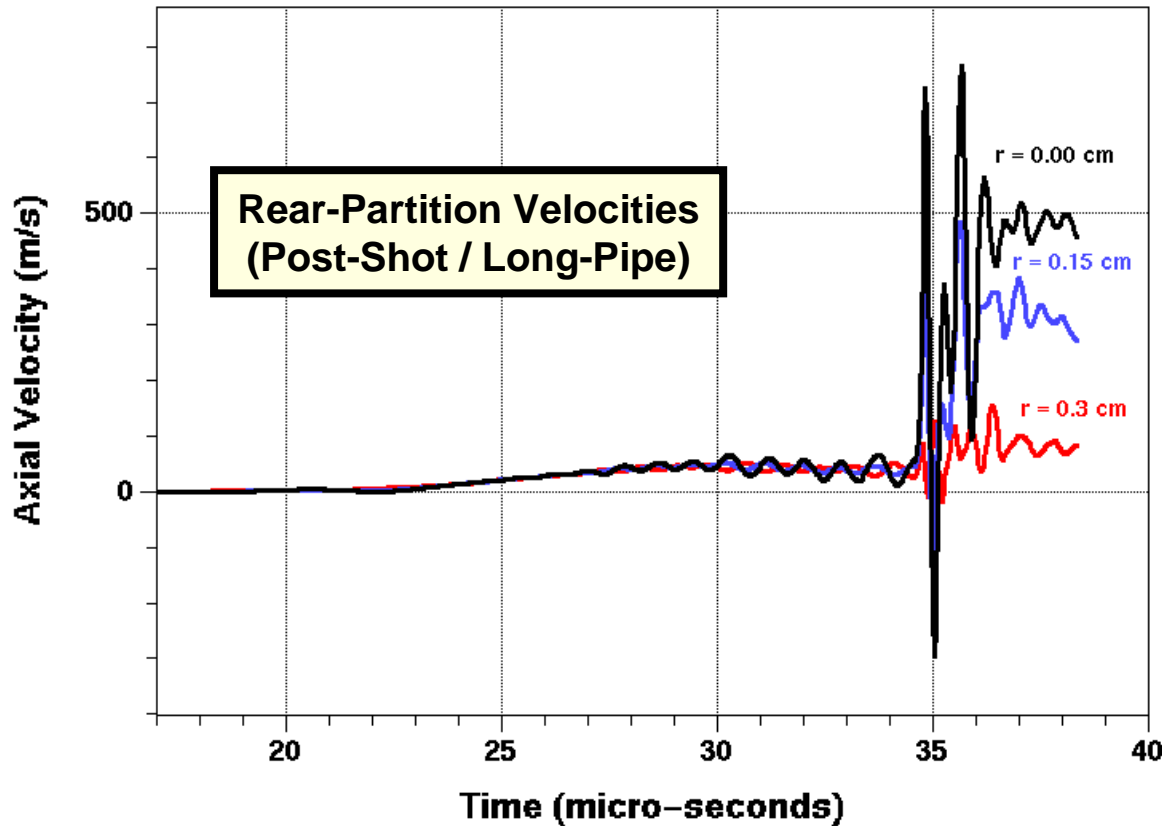


- *These calculated pressure histories are for points half-way through the thickness of the pipe wall at the indicated positions.*
- *The walls and inner partitions were modeled as elastic-plastic materials.*
- *Note the effects of the debris impacts on the traces for partitions 2 and 3.*
- *The midpoint traces do not show these debris-impact signals.*



Velocities for the rear barrier show the effects of both the stress waves in the pipe wall and the debris.

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- *The signals from the stress waves in the pipe wall preceding the debris can be seen at both ~ 20 and $\sim 25 \mu\text{s}$; the debris impact on this barrier occurs at $\sim 34 \mu\text{s}$.*
- *The calculated rear-surface velocities are relatively low, with a maximum value a little over 0.7 km/s ; they are still strongly peaked toward the axis.*

Comparison between experiment and calculation gives only qualitative agreement.

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Long Pipe

Black = Experiment

Red = Calculation

Partition ↓	Arrival Time (μ s)		Partition Vel. (km/s)		
	Toe	Main	VISAR Precursor	VISAR Main	To next partition
Front On-Axis		0.15		3.9	2.95 3.94
Front Off-Axis	0.06	0.15 0.15	0.07	2.3 3.9	2.88 3.78
Middle On-Axis	12.38	17.09 13.0	0.035	0.45 1.7	2.40
Middle Off-Axis	9.33	17.50 13.5	0.076	0.45 0.2	6.3? 2.40
Rear On- Axis	Not Meas. 22.5	Not Meas. 36	Not Meas. 0.04	Not Meas. 0.48	N. A.
Rear Off-Axis	21.16 22.5	~25? 36	~0.01 0.04	>0.1 0.08	N. A.

- Times are shifted to superimpose front-barrier peak particle velocity arrival times.
- Calculations suggest debris is strongly peaked on axis; middle door velocity measurements suggest asymmetry.
- Sectioned short-tunnel shows remaining “door stubs” that are consistent with broad range of debris velocities across tunnel diameter.





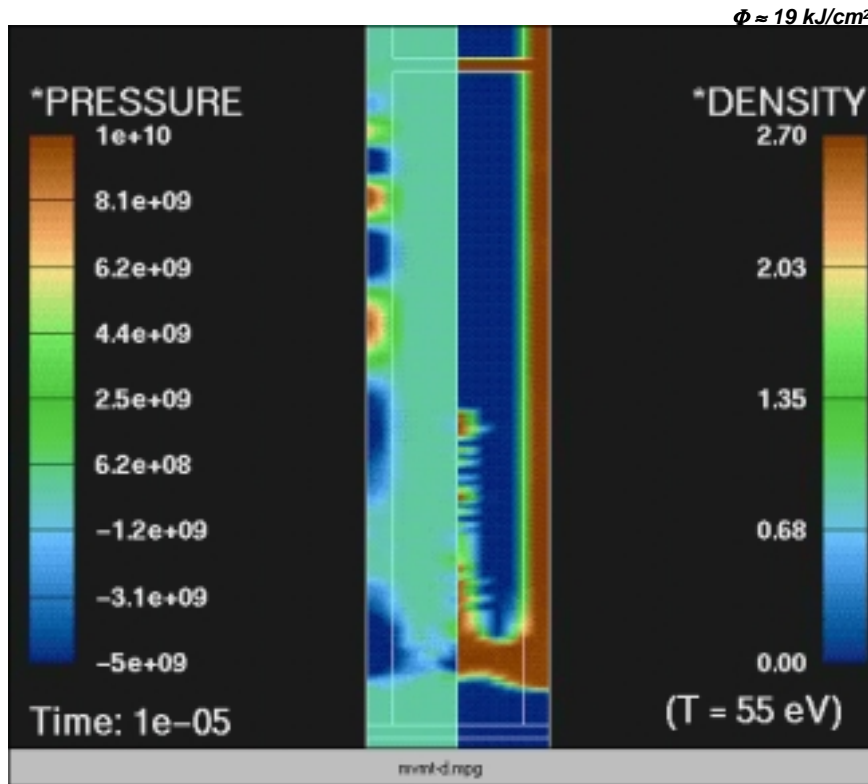
Source temperatures of 100 eV and 55 eV show qualitatively different response phenomenology.

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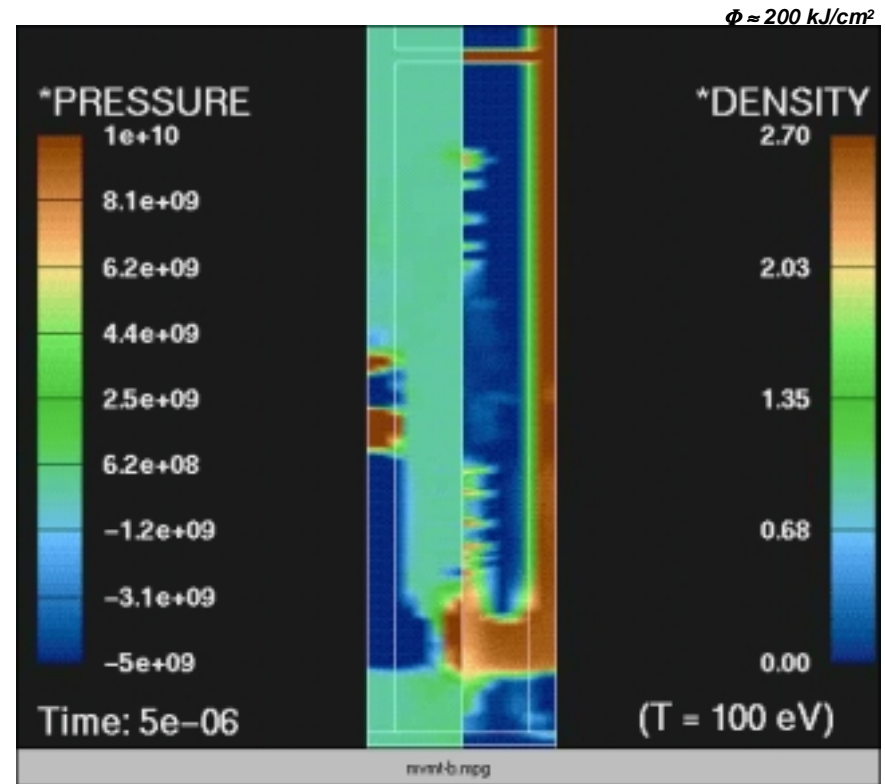


The debris consists of only low-velocity, high-density material -- the second barrier is penetrated at $\sim 21 \mu\text{s}$. The pressure waves precede the debris-generated signals by nearly a factor of two.

Note high-velocity, low-density debris cloud followed by lower-velocity, higher-density material -- the former penetrates the second barrier at $\sim 6 \mu\text{s}$, well ahead of any direct pressure waves in the pipe wall.



$T_{\text{max}} = 55 \text{ eV}; \quad \Phi_{55} \approx \Phi_{100}/11;$
40-ns-wide triangular pulse.



$T_{\text{max}} = 100 \text{ eV}; \quad \Phi_{100} \approx 11 \Phi_{55};$
40-ns-wide triangular pulse.





Overall conclusions . . .

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- **For the radiation-driven jet problem:**
 - > Validation for the ALEGRA models has been achieved through comparison with the NOVA experiment and other calculations.
 - > Experiments scaled up by an order of magnitude on Z show qualitatively similar dynamic response phenomenology.
 - > Diagnostics with ZBL are feasible for the scaled-up experiments.
- **For the partitioned pipe experiment:**
 - > Qualitatively different response phenomenology results from variations in loading conditions.
 - > Measured velocity histories were not well matched, but discrepancies may be due to uncertainties in radiation environment.
 - > ALEGRA was used in a predictive mode to help with experimental design, and new applied capabilities for both Z and ALEGRA have been demonstrated.
- **The ASCI code ALEGRA is under continuing development, but is being used routinely to address practical issues associated with experimental design and applied system-level response.**

